

1 **Formation of mixed paragenesis diamonds during**
2 **multistage growth – constraints from *in situ* $\delta^{13}\text{C}$ – $\delta^{15}\text{N}$ –[N]**
3 **analyses of Koidu diamonds**

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Abstract

20 Inclusion-bearing diamonds from the Koidu kimberlite complex, Sierra Leone (West
21 African Craton) were analyzed *in situ* for carbon and nitrogen isotope compositions, nitrogen
22 concentrations and nitrogen aggregation states. In a suite of 105 diamonds, 78% contain eclogitic
23 mineral inclusions, 17% contain peridotitic mineral inclusions, and 5% - an unusually high
24 proportion - contain co-occurring eclogitic and peridotitic mineral inclusions indicating a mixed
25 paragenesis. Major and trace element compositions of mineral inclusions from two mixed
26 paragenesis diamonds (one with omphacite + Mg-chromite, the other with eclogitic garnet +
27 forsteritic olivine) were determined. The presence of positive Eu anomalies in centrally located
28 omphacite and eclogitic garnet inclusions indicates derivation from subducted protoliths, formed
29 as igneous cumulates in lower oceanic crust. Mg-chromite (Cr# 85.5; Mg# 65.2) and olivine (Mg#
30 94.5) inclusions, located in outer portions of the mixed paragenesis diamonds, have compositions
31 indicative of derivation from strongly depleted cratonic peridotites. Given that the olivine Mg# of
32 94.5 is the highest reported to date for the West African Craton, the eclogitic and peridotitic
33 inclusions in these mixed paragenesis diamonds cannot have precipitated during infiltration of
34 peridotitic substrates by eclogite-derived fluids, as the consequent fluid-rock interaction should
35 lead to Mg# lower than that for the original peridotitic diamond substrate. The different origins of
36 eclogitic and peridotitic inclusions could be explained by physical transport of their host diamonds
37 from eclogitic into peridotitic substrates, possibly along high-strain shear zones, before renewed
38 diamond growth.

39 Based on the $\delta^{13}\text{C}$ - $\delta^{15}\text{N}$ systematics of the entire inclusion-bearing diamond suite from
40 Koidu, three major compositional clusters are identified. Cluster 1 (eclogitic diamond cores; $\delta^{13}\text{C}$
41 = -33.2 to -14.4 ‰ and $\delta^{15}\text{N}$ = -5.3 to +10.1 ‰) bears the isotopic signature of recycled crustal

42 material (\pm a mantle component). Cluster 2 (peridotitic diamonds and including the core of a
43 diamond containing omphacite + Mg-chromite; $\delta^{13}\text{C} = -6.0$ to -1.1 ‰ and $\delta^{15}\text{N} = -4.2$ to $+9.7$ ‰)
44 likely involves mixing of carbon and nitrogen from subducted and mantle sources. Cluster 3 (rims
45 of eclogitic diamonds and including the eclogitic garnet + olivine included diamond and the rim
46 of the omphacite + Mg-chromite included diamond; $\delta^{13}\text{C} = -7.8$ to -3.6 ‰ and $\delta^{15}\text{N} = -7.9$ to -2.1
47 ‰) matches convecting mantle-derived fluids/melts. The distinct isotopic signatures of the three
48 diamond clusters, together with differences in nitrogen aggregation and cathodoluminescence
49 response between diamond cores and rims, suggest episodic diamond growth during multiple
50 fluid/melt pulses.

51

52 **Keywords:** Diamond; carbon isotopes; nitrogen isotopes; mixed paragenesis; West African Craton

53

54 **1. Introduction**

55 Mineral inclusions encapsulated within diamonds preserve pristine information about the
56 conditions of the mantle at the time of diamond formation, as they cannot re-equilibrate with their
57 surroundings during mantle residence or kimberlite eruption. Based on inclusion mineralogy and
58 composition, a first order division can be drawn between common lithospheric diamonds that form
59 within the thick lithospheric mantle keels underpinning long-term stable continental areas (cratons)
60 and rare superdeep diamonds originating from beneath (> 250 km depth) (Stachel and Harris,
61 2008). Lithospheric diamonds are divided into peridotitic, eclogitic and websteritic suites, which
62 represent the mantle substrates for their crystallization. Common peridotitic mineral inclusions are
63 Cr-pyrope garnet, Cr-diopside, forsterite, enstatite, Mg-chromite and Ni-rich sulphide. The

64 eclogitic suite includes, but is not limited to, pyrope-almandine-grossular garnet, omphacite,
65 kyanite, rutile, coesite, corundum and Ni-poor sulphide. The websteritic suite contains minerals
66 that have compositions that are transitional between the peridotitic and eclogitic suites (Meyer,
67 1987; Taylor and Anand, 2004; Stachel and Harris, 2008).

68 Any given lithospheric diamond crystal typically hosts mineral inclusions that are derived
69 from a single suite only (Stachel and Harris, 2008). On some occasions, diamond formation occurs
70 in compositionally changing environments caused by the interaction of metasomatic melts with
71 diamond substrates, which may result in eclogitic and websteritic inclusions being trapped in the
72 same diamond at different stages during diamond growth (Aulbach et al., 2002; Davies et al.,
73 2004a). Diamonds containing both peridotitic and eclogitic mineral inclusions (referred to as
74 “mixed paragenesis” hereafter) are rare and have only been documented from a few localities, e.g.,
75 West Africa (Prinz et al., 1975), Argyle in Western Australia (Hall and Smith, 1984), Monastery
76 in South Africa (Moore and Gurney, 1986), Sloan in USA (Otter and Gurney, 1986), Shengli 1
77 and Pipe 50 in China (Wang, 1998) and the Lac de Gras area in Canada (Davies et al., 2004b).
78 Diamonds of mixed paragenesis are thought to document multiple diamond growth events that
79 occurred in different substrates (Wang, 1998) although other origins are possible, such as
80 modification of mineral compositions by melt-rock reaction in the mantle, accompanied by
81 diamond growth (Mikhail et al., 2021).

82 In this study, we identify five diamonds from the Koidu kimberlite complex in Sierra Leone
83 that contain both eclogitic and peridotitic mineral inclusions – an unusually high number of
84 diamonds relative to the majority of those recovered. We investigate their formation by imaging
85 diamond growth textures, assessing the chemical compositions of the mixed paragenesis mineral
86 inclusions and by comparing the N concentrations, N aggregation states and stable isotope

87 compositions of these unusual diamonds to single-paragenesis diamonds from the same locality.
88 This is the first study to employ multi-collector secondary ion mass spectrometry (SIMS) to
89 measure variations in $\delta^{13}\text{C}$, $\delta^{15}\text{N}$ and N concentration in different growth zones of Koidu diamonds
90 (including peridotitic, eclogitic and mixed paragenesis diamonds) at high spatial resolution,
91 enabling detection of changing growth conditions and therefore the dynamic environment of
92 diamond formation.

93

94 **2. The West African Craton and Sierra Leone diamonds**

95 Inclusion-bearing diamonds for this study were recovered from the Koidu kimberlite
96 complex and an adjacent sedimentary basin, in the Kono District of eastern Sierra Leone, situated
97 on the Archean Man Shield in the southern part of the West African Craton. The Archean rocks in
98 the Man Shield comprise mainly of 3.26–2.85 Ga tonalite-trondhjemite-granodiorite (TTG)
99 gneisses (up to 3.6–3.5 Ga in age), supracrustal belts containing a basalt-komatiite sequence
100 overlain by sediments, and ca. 2.8 Ga granitoids formed in a craton-wide thermal event (Rollinson,
101 2016). The whole rock Re-Os isochron age of low-MgO eclogite xenoliths from the Koidu
102 kimberlite complex was found to be 3.4 ± 0.8 Ga, indicating their formation during Archean
103 subduction (Barth et al., 2002a). Recent Sr-U-Pb dating of clinopyroxenes from Koidu eclogite
104 xenoliths also suggested at least Proterozoic and possibly Archean ages, with a preferred age of
105 2.7 Ga (Aulbach et al., 2019a). The Neoproterozoic break-up of Rodinia may have led to intrusion
106 of a kimberlite-like metasomatic agent in the West African lithosphere (Aulbach et al., 2019b).
107 This metasomatic event may have prompted diamond formation at ca. 0.6 Ga, as indicated by Re-
108 Os isochron ages of sulphide inclusions in diamonds from Zimmi, West Africa, with radiogenic
109 and sulphur stable isotope compositions requiring a subduction origin (Smit et al., 2016, 2019a).

110 The Koidu kimberlite complex is one of a few localities worldwide where eclogites
111 constitute the only type of mantle xenoliths (Tompkins and Haggerty, 1984; Hills and Haggerty,
112 1989; Fung and Haggerty, 1995). Despite the absence of peridotite xenoliths, indicator minerals
113 (heavy media separates) from the Koidu kimberlites are both peridotitic and eclogitic (Skinner et
114 al., 2004). A recent study of Koidu indicator minerals documented a large proportion of garnets
115 with high Cr and low Ca contents, indicating the presence of strongly depleted harzburgites or
116 dunites in the underlying lithospheric mantle (Harder et al., 2013).

117 Diamonds containing either eclogitic or peridotitic inclusions have been recovered from
118 Sierra Leone, though only very limited information has been published. Recovered inclusions
119 include olivine, enstatite, Cr-pyrope, chromite (Meyer and Boyd, 1972) and sulphide (Deines and
120 Harris, 1995; Smit et al., 2016). Carbon isotope analyses of Sierra Leone diamonds are also scarce,
121 with only a few chromite- and sulphide-bearing diamonds from Koidu and Zimmi (Deines and
122 Harris, 1995; Smit et al., 2019b, respectively) having been analyzed, and no N isotope
123 compositions reported.

124

125 **3. Sample description**

126 A total of 111 inclusion-bearing rough diamonds 2–4 mm in size were studied. The total
127 weight of the diamonds is 10.7 carats. All diamonds in the sample suite are colourless except for
128 one pale yellow and one light brown diamond. Their morphology is dominated by octahedra (n =
129 85), followed by rounded dodecahedra (n = 11), transitional octahedra-dodecahedra (n = 7),
130 irregular shapes (n = 4), macles (n = 2), an aggregate (n = 1) and an intergrowth of two octahedra
131 (n = 1). Descriptions of diamonds are summarized in Table S1.

132 Resorption features are observed on all diamonds, with negative trigons occurring on all
133 octahedral faces and aligning in parallel rows where plastic deformation lines are present. Other
134 common resorption features include shield-shaped laminae and hexagonal pits on octahedral faces,
135 hillocks, corrosion sculptures and micro-disk patterns on rounded dodecahedral faces, terraces
136 around the three-fold axes of dodecahedral crystals, and tetragons on presumed {100} faces of
137 irregular-shaped diamonds. Surface textures not restricted to specific crystal faces include
138 deformation lines, ruts, fractures, and a green irradiation spot on an octahedral diamond. Voids of
139 cubo-octahedral shapes are found on the surface of some diamonds, likely representing cavities
140 previously occupied by mineral inclusions. Examples of common surface textures of Koidu
141 diamonds can be found in Fig. S1.

142 Six diamonds contained only crustal mineral inclusions (epigenetic) and thus they were not
143 further investigated. From 105 diamonds, 370 primary mineral inclusions were recovered, and we
144 found that 82 diamonds (78%) contained eclogitic inclusions, 18 diamonds (17%) contained
145 peridotitic inclusions, and five diamonds (5%) contained both eclogitic and peridotitic inclusions.
146 The chemical composition of mineral inclusions in single-paragenesis Koidu diamonds will not be
147 discussed further in the current manuscript.

148 The five diamonds of mixed paragenesis are 130-9 (omphacite + olivine), 130-31 (coesite
149 + olivine), 133-6 (omphacite + Mg-chromite), 138-7 (eclogitic garnet + olivine) and 146-1 (olivine
150 + Mg-chromite + coesite). Omphacite, olivine and coesite inclusions are all colourless, but were
151 identified using their chemical compositions. The spatial distribution of particular inclusions
152 within their host diamonds and consequently the direction of changes in inclusion composition
153 could only be investigated for two diamonds containing mineral inclusions with clearly
154 distinguishable colours — diamonds 133-6 and 138-7. The chemical composition of the mineral

155 inclusions in these two diamonds are discussed in detail below, whereas the inclusion compositions
156 of the other mixed paragenesis diamonds are reported in Table S3. In both diamonds, eclogitic
157 inclusions are located in the centre and peridotitic inclusions near the rim (Fig. 1).

158

159 **4. Analytical methods**

160 **4.1 Mineral inclusions**

161 Mineral inclusions were mounted with epoxy resin in 6 mm diameter brass tubes and
162 polished. Initial mineral identification was achieved by back-scattered electron (BSE) imaging and
163 energy dispersive spectroscopy (EDS). Major and minor element compositions were determined
164 using CAMECA SX100 and JEOL JXA-8900R electron probe microanalyzers (EPMA), both
165 equipped with five wavelength-dispersive spectrometers (WDS). The instruments were operated
166 at an accelerating voltage of 20 kV, a beam current of 20 nA, and a fully-focussed beam with a
167 diameter < 1 μm . Reference materials include metals and natural and synthetic minerals. For all
168 elements, the $K\alpha$ emission lines were employed for analysis. The counting time was 20–100 s on
169 the peak, with resulting oxide detection limits typically ≤ 0.02 wt%. Three spots were measured
170 on each grain and, after assessing compositional homogeneity, the compositions of the spots were
171 averaged.

172 Trace elements (including REE, Ti, V, Ni, Rb, Sr, Zr, Nb, Ba and Hf) in garnet and
173 clinopyroxene inclusions were determined using a Resonetics M-50-LR 193 nm ArF excimer laser
174 ablation system, with a Laurin-Technic S-155 two-volume ablation cell, coupled with a Thermo
175 Scientific Element IIXR inductively coupled plasma mass spectrometer (LA-ICP-MS). Mineral
176 inclusions were ablated with a spot size of 23–90 μm at a frequency of 10 Hz and a laser fluence

177 of $\sim 4 \text{ J/cm}^2$. Analysis of each sample includes 40 s background collection followed by 60 s sample
178 ablation/measurement and 40 s washout. Detection limits are typically $\leq 40 \text{ ppb}$ for REE, V, Rb,
179 Sr, Zr, Nb, Ba and Hf, and $\leq 1 \text{ ppm}$ for Ti and Ni. Calibration was achieved using the NIST SRM
180 612 glass standard and ^{43}Ca as internal standard for data processing. USGS reference glass BIR-
181 1G was analysed as an unknown and the results were compared to the GeoRem preferred values
182 (Jochum et al., 2005) to assess repeatability and accuracy.

183 **4.2 Host diamonds**

184 Infrared absorption spectra of diamonds were obtained before crushing, using a Thermo
185 Fisher Nexus 470 Fourier transform infrared (FTIR) spectrometer equipped with a Continuum IR
186 microscope. An aperture size of $100 \times 100 \text{ }\mu\text{m}$ was used in transmission mode. Spectra were
187 acquired by averaging 200 scans at a spectral resolution of 1 cm^{-1} with a spectral range of 4000--
188 650 cm^{-1} . Baseline-corrected spectra were normalized to 1 cm diamond thickness, where the
189 absorption coefficient for the intrinsic absorption of diamond at 1995 cm^{-1} is approximately 11.94
190 cm^{-1} (e.g., Howell et al., 2012). Nitrogen concentrations and aggregation states of diamonds were
191 calculated by deconvolution of the normalized spectra, where $[\text{N}_A] = 16.5 \times \mu\text{A}$ (absorption
192 coefficient at 1282 cm^{-1} for the A-centre; Boyd et al., 1994a) and $[\text{N}_B] = 79.4 \times \mu\text{B}$ (absorption
193 coefficient at 1282 cm^{-1} for the B-centre; Boyd et al., 1995). The detection limit for N is $\sim 5 \text{ at.ppm}$.
194 More than one spectrum was collected for diamonds at different spots. Note that the N distribution
195 in diamond may be heterogeneous, and the infrared absorption spectrum is an integrated signal
196 collected over the optical path of the diamond, which may contain different growth zones with
197 distinctly different N contents.

198 Diamond fragments were first cast in epoxy and then ground and polished. Subsequently,
199 the fragments were co-mounted in indium with a SIMS reference material – diamond S0270. Prior

200 to analysis, cathodoluminescence (CL) images of diamond fragments were obtained using a Zeiss
201 EVO MA15 scanning electron microscope (SEM) equipped with a parabolic mirror coupled to a
202 high sensitivity broadband photomultiplier tube to reveal the diamond internal growth textures.
203 The mount was coated with Au to prevent charging during SEM operation. The scanning electron
204 microscope was operated at a voltage of 15 kV and a beam current of 3–5 nA. Subsequently, C
205 and N isotope compositions and N abundances of the mounted diamond fragments were
206 determined using a CAMECA IMS-1280 multi-collector ion microprobe. Two to eight
207 measurement spots were set on each diamond fragment to cover all growth zones. The primary
208 beam of 20 keV $^{133}\text{Cs}^+$ ions was focused to a beam diameter of approximately 10 μm , with beam
209 currents of 2 nA, 2.5–3.0 nA and 0.6 nA for the analyses of C isotope composition, N isotope
210 composition and N abundance, respectively. Carbon isotopes were analyzed first, followed by N
211 abundance and N isotope measurements from the same spot location. The detection limit for N
212 abundance is ~ 0.1 at.ppm. Nitrogen isotope compositions were only measured on spots with N
213 abundance > 65 at.ppm. Carbon isotope compositions are reported as $\delta^{13}\text{C}_{\text{VPDB}}$, which is the
214 normalized difference of the $^{13}\text{C}/^{12}\text{C}$ ratio of the sample relative to that of the Vienna Pee Dee
215 Belemnite standard ($^{13}\text{C}/^{12}\text{C}_{\text{VPDB}} = 0.01118$; Coplen et al., 2002). Nitrogen isotope compositions
216 are reported as $\delta^{15}\text{N}_{\text{AIR}}$, which is the normalized difference of the $^{15}\text{N}/^{14}\text{N}$ ratio of the sample
217 relative to that of the atmosphere ($^{15}\text{N}/^{14}\text{N}_{\text{AIR}} = 0.003677$; Junk and Svec, 1958). The analytical
218 sequences for C and N isotopes interspersed measurements of unknowns with diamond reference
219 S0270 ($\delta^{13}\text{C}_{\text{VPDB}} = -8.88 \pm 0.10$ ‰; $\delta^{15}\text{N}_{\text{AIR}} = -0.40 \pm 0.50$ ‰; Stern et al., 2014) in a 4:1 ratio.
220 Uncertainties of individual $\delta^{13}\text{C}_{\text{VPDB}}$ analyses for diamond S0270 and unknowns are typically
221 ± 0.14 ‰ (2σ). Uncertainties of individual $\delta^{15}\text{N}_{\text{AIR}}$ analyses are typically ± 0.60 ‰ (2σ) for diamond
222 S0270, and typically range from ± 0.50 to ± 4.0 ‰ (2σ) for unknowns with N concentration ranging

223 from 2500 to 50 at.ppm, respectively. Additional details of the analytical procedure were outlined
224 in Stern et al. (2014).

225

226 **5. Results**

227 **5.1 Mineral inclusions from mixed paragenesis diamonds**

228 **5.1.1 Diamond 133-6**

229 Diamond 133-6 contains four omphacites (clinopyroxenes with a jadeite component ≥ 20
230 and < 80 mol%; Clark and Papike, 1968) within a radius of ~ 500 μm in the centre and one Mg-
231 chromite in the rim (Fig. 1a). The omphacites were surrounded by small fractures, which did not
232 extend to the surface of the diamond. Backscattered electron images indicated that all inclusions
233 have homogeneous compositions throughout the grains. Inter-grain variation among the four
234 omphacites was negligible, thus an average major element composition is reported.

235 The omphacites have high molar Mg# ($100 \times \text{Mg}/[\text{Mg}+\text{Fe}] = 82.2$) and jadeite component
236 ($100 \times 2\text{Na}/[2\text{Na}+\text{Ca}+\text{Mg}+\text{Fe}] = 63.9$ mol%). The Ca-Mg-Fe compositions of these omphacites are
237 similar to the omphacites from Koidu gabbroic eclogites (classified based on bulk-rock $[\text{Eu}/\text{Eu}^*]_{\text{N}}$
238 ≥ 1.05 , where $\text{Eu}^* = [\text{Sm}_{\text{N}} \times \text{Gd}_{\text{N}}]^{0.5}$; Aulbach et al., 2019b) and low-MgO eclogites (bulk-rock
239 MgO content = 6–13 wt%; Hills and Haggerty, 1989; Aulbach et al., 2019b) (Fig. 2). The average
240 Cr_2O_3 content in these omphacites is very low (≤ 0.02 wt%) (Table 1).

241 Trace element compositions were determined for three of the omphacites from diamond
242 133-6 (Table 2). They have very similar REE_{N} patterns and show positive Eu anomalies (defined
243 as $[\text{Eu}/\text{Eu}^*]_{\text{N}} > 1$), with $[\text{Eu}/\text{Eu}^*]_{\text{N}} = 1.63$ – 1.78 . The REE_{N} patterns of these omphacites are clearly

244 different from omphacites in Koidu eclogite xenoliths (Aulbach et al., 2019b) (Fig. 3). The
245 omphacite inclusions are enriched in LREE relative to CI-chondrite, but have unusually low
246 MREE and HREE compared to clinopyroxene inclusions in diamonds worldwide (Stachel and
247 Harris, 2008), resulting in high LREE/HREE ratios. The concentrations of the HREE Tm, Yb and
248 Lu are below the limit of detection, except for Yb in one omphacite. They also are enriched in Nb,
249 Sr, Ti and V relative to CI-chondrite (Fig. 4).

250 The Mg-chromite inclusion has a cubo-octahedral morphology and no associated fractures,
251 implying a synchronous relationship with the growth zone of the host diamond in which it occurs.
252 It has high Mg# (65.2) and Cr# ($100 \times \text{Cr} / [\text{Cr} + \text{Al}] = 85.5$), typical for peridotitic spinel inclusions
253 in diamond worldwide (Stachel and Harris, 2008).

254 **5.1.2 Diamond 138-7**

255 Diamond 138-7 contains five eclogitic garnets in the centre and one olivine in the rim (Fig.
256 1b). No fractures are observed around the inclusions nor elsewhere in this diamond. All inclusions
257 grains are internally compositionally homogeneous, without any sign of alteration.

258 Garnets show significant grain to grain compositional variations in SiO₂ (39.5–41.9 wt%),
259 Al₂O₃ (21.9–23.9 wt%), MgO (8.1–9.7 wt%) and CaO (16.7–19.3 wt%) (Table 1). The very high
260 CaO contents are unusual and worldwide, garnet inclusions with CaO contents ≥ 17 wt% have
261 been reported only from a few localities, most prominently Klipspringer in South Africa
262 (Westerlund and Gurney, 2004), the New South Wales alluvials (Davies et al., 2003) and Argyle
263 in Australia (Jaques et al., 1989), but have not been observed at other localities on the West African
264 Craton. Among Koidu eclogite xenoliths, similar high CaO in garnet is only seen in a few low-
265 MgO kyanite or corundum eclogites (Hills and Haggerty, 1989) (Fig. 5).

266 Trace element compositions of three garnets from diamond 138-7 were determined (Table
267 2). Their REE_N patterns are very similar to one another, with small differences in La and from Eu
268 to Lu (Fig. 6). All garnets have positive Eu anomalies, with [Eu/Eu*]_N = 1.10–1.41. They have
269 REE_N patterns typical of cratonic eclogite xenoliths, characterised by increasing abundances from
270 subchondritic La to about 10x chondritic MREE and HREE (Sm to Lu). Their MREE to HREE
271 contents overlap with the field of garnets from Koidu gabbroic eclogites (Aulbach et al., 2019b),
272 with the LREE La to Pr above the xenolith garnet range. Strontium and Nb abundances are
273 subchondritic (Fig. 7). There are positive correlations between trace element and major element
274 compositions, where Nb and Ti increase with CaO abundance in these garnets (Tables 1 and 2).

275 The single olivine inclusion is elongated in shape. It has a very high Mg# (94.5),
276 documenting the strongly depleted nature of the harzburgitic to dunitic diamond substrates. On the
277 West African Craton, such high Mg# have only been reported for retrograde olivines occurring as
278 inclusions in sublithospheric diamonds (Kankan, Guinea; after primary wadsleyite or ringwoodite;
279 Stachel et al., 2000a), but not for olivine in lithospheric diamonds (Meyer and Boyd, 1972 ; Stachel
280 and Harris, 1997; Stachel et al., 2000b).

281 **5.2 $\delta^{13}\text{C}$, $\delta^{15}\text{N}$ and N concentrations in Koidu diamonds**

282 **5.2.1 Peridotitic diamonds**

283 Peridotitic diamonds from Koidu (18 diamonds; 54 SIMS analyses; Tables 3 and S3) have
284 a restricted range in $\delta^{13}\text{C}$ values from -6.0 to -1.1 ‰, well within the range of peridotitic diamond
285 worldwide (Stachel et al., 2009) (Fig. 8a). Their N abundances (all values stated in section 5.2
286 were obtained by SIMS) and $\delta^{15}\text{N}$ values range from 0.4 to 920 at.ppm (median = 80 at.ppm) and
287 -4.2 to +9.7 ‰, respectively, and define one of three clusters (Cluster 2) shown in Fig. 9 (see

288 Section 6.4 for detailed discussion). Three peridotitic diamonds have mantle-like isotopic
289 signatures (defined as $\delta^{13}\text{C} = -5 \pm 3 \text{‰}$ and $\delta^{15}\text{N} = -5 \pm 3 \text{‰}$; Cartigny et al., 2014) with $\delta^{13}\text{C}$ values
290 of -5.7 to -2.6 ‰ and $\delta^{15}\text{N}$ values of -4.2 to -2.1 ‰ (Fig. 9), accompanied by variable N contents
291 (1–680 at.ppm). Most diamonds have small internal variations in C isotope composition (average
292 $\sigma < 0.2 \text{‰}$), except for four diamonds (130-21, 130-25, 134-2 and 136-1; Table S4) which have
293 intra-diamond $\delta^{13}\text{C}$ variations up to 2.8 ‰.

294 **5.2.2 Eclogitic diamonds**

295 Koidu eclogitic diamonds in this study (82 diamonds, 321 SIMS analyses; Tables 3 and S3)
296 show highly variable isotope compositions and N concentrations, indicating growth during
297 multiple events. Based on their $\delta^{13}\text{C}$ and $\delta^{15}\text{N}$ values, they define two additional clusters (Clusters
298 1 and 3; Fig. 9), which will be discussed in Section 6.4 in detail.

299 Based solely on their $\delta^{13}\text{C}$ values, the eclogitic diamonds can be separated into three
300 subgroups:

301 *^{13}C -enriched diamond:* This group contains only a single diamond with the highest $\delta^{13}\text{C}$
302 values among Koidu eclogitic diamonds. The bulk of this diamond has $\delta^{13}\text{C}$ values from -1.8 to
303 0.0 ‰ and negligible N content. Lower $\delta^{13}\text{C}$ values (-4.2 ‰) are observed in a more N-rich (50–
304 80 at.ppm) growth zone. $\delta^{15}\text{N}$ values were not determined for this overall N-poor diamond.

305 *^{13}C -depleted diamond:* Diamonds in this group have $\delta^{13}\text{C}$ values from -33.2 to -15.1 ‰,
306 with ~80% of the data falling in the range of -30 to -25 ‰ (Fig. 8b). Compared to Koidu peridotitic
307 diamonds, this group has a broader range of N abundances (0.4 to 2080 at.ppm; median = 30
308 at.ppm) and $\delta^{15}\text{N}$ values (-5.3 to +9.9 ‰; Fig. 9). Most diamonds are internally homogeneous in
309 carbon isotope composition, but with a spread in $\delta^{15}\text{N}$ values of up to 5.4 ‰. Six diamonds in this

310 group have a less ^{13}C -depleted outermost growth zone ($\delta^{13}\text{C} = -19.2$ to -15.1 ‰) compared to the
311 core ($\delta^{13}\text{C} = -29.6$ to -24.9 ‰).

312 *Diamonds with cores depleted in ^{13}C and rims with mantle-like C isotopic signature:*

313 Twenty-four diamonds have a clear core-rim structure (core and rim have distinctly different CL
314 brightness), where the core zone has low $\delta^{13}\text{C}$ values (-30.2 to -14.4 ‰) and generally low N
315 concentrations (1–890 at.ppm; median = 30 at.ppm), while the rim has higher, mantle-like $\delta^{13}\text{C}$
316 values (-7.8 to -5.6 ‰) and high N concentrations (610–2200 at.ppm; median = 870 at.ppm) (Figs.
317 8c and 10). For the core zones, only 12 diamonds have sufficiently high N abundances (> 65 at.ppm)
318 to allow for precise N isotope analysis; $\delta^{15}\text{N}$ values in the cores range from -5.6 to $+8.1$ ‰,
319 overlapping with the range for the subgroup of ^{13}C -depleted eclogitic diamonds (see above). The
320 high N abundances in the rims allow the measurement of N isotope compositions for all diamonds
321 in this group, indicating mantle-like $\delta^{15}\text{N}$ values between -7.9 and -2.6 ‰ (Fig. 9). One diamond
322 in this group has a distinct intermediate layer between the core and the rim, with an average $\delta^{13}\text{C}$
323 value of -15.3 ‰ that is ~ 10 ‰ higher than the core zone and ~ 8 ‰ lower than the rim, but similar
324 to the outermost growth zones of the six zoned ^{13}C -depleted eclogitic diamonds (see above). $\delta^{15}\text{N}$
325 values of this diamond fluctuate from -0.5 ‰ in the core zone to $+10.1$ ‰ in the intermediate layer
326 to -5.5 ‰ in the rim.

327 **5.2.3 Mixed paragenesis diamonds**

328 The isotope compositions of all studied mixed paragenesis diamonds (5 diamonds, 89
329 SIMS analyses) are shown in Figs. 8d and 9, and together with N concentrations listed in Tables 3
330 and S3. These mixed paragenesis diamonds lie within the three clusters (Fig. 9) mentioned above
331 and will be discussed in detail in Section 6.4.

332 Diamond 133-6 (omphacite and Mg-chromite included) is comprised of inner growth zones
333 and a rim with similar $\delta^{13}\text{C}$ (-4.8 to -3.6 ‰) but distinct $\delta^{15}\text{N}$ values and N contents. The inner
334 growth zones have a broad range of N isotope compositions ($\delta^{15}\text{N} = -6.2$ to $+4.8$ ‰) and low N
335 contents (1–380 at.ppm). The rim has a restricted, mantle-like N isotopic signature ($\delta^{15}\text{N} = -7.9$ to
336 -6.7 ‰) and high N contents (520–750 at.ppm). Core-to-rim transects show that within the
337 innermost zone of this diamond (fragments 1 and 3 in Fig. 11) a minor outward increase in $\delta^{13}\text{C}$
338 values (from -4.7 to -3.7 ‰) is accompanied by decreasing N concentrations (from 330 to 20
339 at.ppm), but no trend in $\delta^{13}\text{C}$ values is observed for the intermediate or rim zones.

340 Diamond 138-7 (eclogitic garnet and olivine included) does not have a discernible core-
341 rim zonation. Carbon and N isotope compositions vary ($\delta^{13}\text{C} = -5.3$ to -4.1 ‰; $\delta^{15}\text{N} = -9.8$ to -5.3
342 ‰) within the mantle range. Nitrogen abundance varies widely, from 1 to 1050 at.ppm. Although
343 multiple growth zones and some mild resorption at the outer boundary of the inner growth zone
344 are observed, the zones generally have a homogeneous CL response and transects across them do
345 not reveal coherent trends in $\delta^{13}\text{C}$ values.

346 Diamonds 130-31 (coesite and olivine included) and 146-1 (coesite, Mg-chromite and
347 olivine included) have isotope compositions ($\delta^{13}\text{C} = -4.7$ to -1.3 ‰; $\delta^{15}\text{N} = -2.3$ to $+5.8$ ‰) and N
348 contents (0.3–340 at.ppm) within the range of peridotitic Koidu diamonds. Diamond 130-9
349 (omphacite and olivine included) has an isotopic signature ($\delta^{13}\text{C} = -24.8$ to -22.6 ‰; $\delta^{15}\text{N} = -4.1$
350 ‰) and N contents (0.9–180 at.ppm) similar to the ^{13}C -depleted eclogitic diamonds in this study.

351 **5.3 Nitrogen aggregation in episodically grown diamonds**

352 Nitrogen is incorporated into the diamond lattice in the form of single substitutional N (C-
353 centres; Taylor et al., 1996; Type Ib) during crystallization. Residence in the mantle at high

354 temperature leads to the diffusion and combination of single N atoms into pairs (A-centres; Davies,
355 1976; Type IaA). Nitrogen pairs further aggregate to a structure of four N atoms surrounding a
356 vacancy (B-centre; Jones et al., 1992; Type IaB), along with the generation of two side products –
357 the N3 centre (three N atoms surrounding a vacancy) and platelets (aggregations of interstitial C
358 atoms) (Woods, 1986). The rate of N aggregation depends on three major factors – mantle
359 residence temperature, mantle residence time and N concentration in diamonds (e.g., Taylor et al.,
360 1990).

361 Cathodoluminescence images of the episodically grown diamonds show a relative
362 relationship of bright cores and dark rims (Fig. 10), indicating low concentrations of optically
363 active lattice defects to cause CL in the rims. Infrared absorption spectra (Table S1, S2 and Fig.
364 12) show that the brighter cores contain N in higher aggregation states (Type IaAB with $\geq 30\%$ B),
365 where N3 centres (with a luminescent wavelength at 415 nm; Zaitsev, 2001) are present, while the
366 rims represent overgrowths of diamond with lower N aggregation states dominated by A-centres,
367 which are known to quench luminescence (Vasil'ev et al., 2004). Enhanced N aggregation states
368 in the cores indicate that they had longer mantle residence times or higher mantle storage
369 temperatures compared to the rims (Taylor et al., 1990; Leahy and Taylor, 1997), documenting
370 that cores and rims in these diamonds formed during distinct growth events separated in time or in
371 temperature.

372

373 **6. Discussion**

374 **6.1 Origin of eclogitic mineral inclusions in Koidu mixed paragenesis diamonds**

375 Positive Eu anomalies in omphacites from diamond 133-6 and garnets from diamond 138-
376 7 indicate accumulation of plagioclase in their protoliths (Schmickler et al., 2004), suggesting a
377 protolith origin as gabbro cumulates. Major element compositions of both omphacite and garnet
378 inclusions are similar to some Koidu kyanite- and corundum-bearing low-MgO eclogite xenoliths
379 (Hills and Haggerty 1989), and trace element compositions (MREE to HREE) of garnet inclusions
380 are similar to garnet from Koidu gabbroic eclogite xenoliths (Aulbach et al., 2019b), suggesting
381 that the eclogitic inclusions in mixed paragenesis diamonds inherited compositional characteristics
382 of their eclogitic diamond substrates.

383 Omphacites from diamond 133-6 have highly fractionated REE_N patterns ($La_N/Yb_N = 422$)
384 that are very different from typical eclogitic clinopyroxenes, which usually have humped patterns
385 with $La_N/Nd_N < 1$ and higher HREE contents (e.g., Jacob, 2004). Omphacites with overall very
386 similar REE_N and HFSE_N patterns at lower LREE (Figs. 3 and 4) were, however, observed in two
387 low-MgO eclogites from Obnazhennaya in Siberia (Sun et al., 2020) and a kyanite eclogite from
388 Bellsbank in South Africa (Shu et al., 2016). This similarity in trace element patterns does not fully
389 extend to major element compositions, where the omphacites from diamond 133-6 and from
390 Obnazhennaya and Bellsbank are distinct in their Mg# (82.2 versus 88.8 and 88.4, respectively)
391 and jadeite content (63.9 versus 22.1 and 46.5 mol%, respectively). The protoliths of
392 Obnazhennaya low-MgO and Bellsbank kyanite eclogites were suggested to be olivine gabbro
393 cumulates (Shu et al., 2016; Sun et al., 2020). Eclogites with cumulate protoliths are inferred to
394 have incorporated a trapped melt component and subsequently, during subduction, to have
395 experienced eclogite-facies melt extraction (Aulbach et al., 2007; Aulbach and Jacob, 2016; Sun
396 et al., 2020). Without coexisting garnet, we cannot evaluate the origin of the clinopyroxene
397 inclusions in diamond 133-6 in detail, but based on their high similarity in trace element

398 composition to the Obnazhennaya and Bellsbank clinopyroxenes we suggest a similar origin, i.e.,
399 that the clinopyroxene inclusions derived from oceanic cumulates (possibly with trapped melt) that
400 were subducted and metamorphosed. In addition, the diamond substrate could have been affected
401 by a subsequent stage of metasomatic LREE enrichment to explain the significantly higher
402 contents in La to Pr relative to Obnazhennaya and Bellsbank clinopyroxenes (Fig. 3).

403 The high and variable CaO contents in the garnets from diamond 138-7 may reflect
404 continued metasomatic enrichment during diamond growth. This was previously suggested for an
405 eclogitic diamond from the Mir kimberlite containing 35 garnet inclusions with a wide range of
406 CaO contents (Sobolev et al., 1998). Multiple olivine inclusions with different CaO contents in a
407 single Kankan diamond were also attributed to metasomatic Ca enrichment during diamond
408 growth (Stachel et al., 2000b). The observed positive correlations of Ca with Nb and Ti for the
409 garnets from diamond 138-7 indicate that the metasomatic process accompanying diamond
410 formation could have affected trace elements. Elevated LREE in the garnet inclusions relative to
411 garnets from Koidu gabbroic eclogite xenoliths (Aulbach et al., 2019b) appear to support this
412 interpretation. A possible metasomatic agent driving diamond precipitation and enrichment in Ca
413 and LREE could be a low-Mg carbonatitic high-density fluid or a carbonatitic melt (Sobolev et al.,
414 1998; Stachel et al., 2000b; Klein-BenDavid et al., 2009). Alternatively, the difference in LREE
415 concentrations between inclusion and gabbroic eclogite xenolith garnets may be attributed to
416 crystal-chemical effects: the Ca# ($\text{Ca}/[\text{Ca}+\text{Mg}+\text{Fe}+\text{Mn}]$) of garnet strongly influences partitioning
417 of trace elements between garnet and pyroxene, with both the REE concentrations and the
418 LREE/HREE ratio in garnet increasing together with $\text{Ca}\#_{\text{garnet}}$ (Harte and Kirkley 1997; Aulbach
419 et al., 2017).

420 Either way, two key observations suggest that the garnet inclusions in diamond 138-7
421 predominantly reflect their substrate composition: (1) the presence of positive Eu anomalies and
422 (2) the overall shape of the REE_N patterns, which are typical for garnets in eclogites with low-
423 pressure protoliths that experienced some cumulate enrichment (lowering the REE overall) and
424 possibly partial melting during or after subduction (preferentially extracting LREE) (Ireland et al.,
425 1994; Barth et al., 2001; Stachel et al., 2004; Aulbach and Jacob, 2016).

426 **6.2 Progressive precipitation of eclogitic and then peridotitic inclusions from an evolving** 427 **eclogite-derived fluid**

428 Geochemical modelling, based on the Extended Deep Earth Water (DEW) model, indicated
429 that the interaction between eclogite-derived fluids and a range of mantle peridotite compositions
430 may result in the progressive formation of eclogitic, websteritic and peridotitic minerals (Mikhail
431 et al., 2021). If captured in the form of inclusions during associated diamond growth, this model
432 predicts the occurrence of inclusions of different paragenesis in single diamonds. The calculated
433 models predict 2-3 vol% spinel in the final mineralogy resulting from fluid-peridotite interactions,
434 which seemingly appears to fit with our observation of spinel (chromite) in mixed paragenesis
435 diamond 133-6. With the caveat that the minor element Cr is not included in the DEW model runs,
436 the predicted spinel-group mineral, however, is magnetite, which is a very rare inclusion in
437 diamond (Stachel et al., 1998). Instead, the spinel inclusion found in the rim of diamond 133-6 is
438 a Mg-chromite, with the high Mg# (65.2) and Cr# (85.5) typical of garnet-facies spinels found in
439 cratonic peridotite xenoliths and as inclusions in diamonds (McDonough and Rudnick, 1998;
440 Stachel and Harris 2008). Similarly, the olivine inclusion near the rim of mixed paragenesis
441 diamond 138-7 bears the highly magnesian character (Mg# 94.5) of strongly depleted cratonic
442 peridotites, very much unlike the expected product of a process that converts depleted peridotite

443 into websterite (Mallik and Dasgupta, 2012, 2013). As discussed in the preceding section, the
444 centrally located eclogitic garnet and omphacite inclusions also are very much alike eclogite
445 xenolith minerals and bear the characteristics of subducted gabbroic protoliths rather than those of
446 an eclogite-derived fluid. The very high CaO content of the eclogitic garnet inclusions precludes
447 equilibration with orthopyroxene and consequently, derivation from websterite, which is the
448 product of the modelled fluid-rock interactions (Mikhail et al., 2021). Thus, we conclude that the
449 peridotitic inclusions in our samples are not the result of extended eclogitic fluid-peridotite
450 reactions but instead must have a separate origin from the “coexisting” eclogitic inclusions.
451 Furthermore, with their distinct CL responses and N concentrations, the core and rim zones of
452 diamond 133-6 (Fig. 11) suggest episodic diamond growth rather than fluid evolution during a
453 single metasomatic event.

454 **6.3 Covariation of $\delta^{13}\text{C}$ values and N abundance**

455 In the published literature, N is generally considered as a compatible element in diamond,
456 incorporated through equilibrium partitioning with the diamond-forming fluid/melt, with $K_{\text{N}}^{\text{diamond-}}$
457 $\text{fluid} \geq 2$ for diamond crystallization from a reduced medium (Thomassot et al., 2007; Smit et al.,
458 2019b), and $K_{\text{N}}^{\text{diamond-fluid}} \geq 4$ for diamond growth from an oxidized medium (Stachel et al., 2009;
459 Smart et al., 2011; Petts et al., 2015). Therefore, progressive diamond formation from a single
460 pulse of fluid/melt should lead to a gradual decrease in N concentration.

461 Assuming precipitation from a single C species in the fluid/melt, diamond showing a trend
462 of outward decreasing N concentration accompanied by increasing or decreasing $\delta^{13}\text{C}$ values
463 reflects diamond formation from an oxidized (CO_2 or CO_3^{2-}) or a reduced medium (CH_4),
464 respectively (Deines, 1980). Thus, the core-to-rim transects showing progressively decreasing N
465 concentrations and increasing $\delta^{13}\text{C}$ values within the innermost growth zone of diamond 133-6

466 may indicate an oxidized diamond-forming medium (Fig. 11). Alternatively, precipitation from a
467 mixed CHO fluid, containing both CH₄ and CO₂ in variable ratios, would also lead to ¹³C
468 enrichment during diamond precipitation under fluid-limited conditions (Stachel et al., 2017). The
469 same covariation is not observed in the intermediate zones and the rim of this diamond (Fig. 11),
470 suggesting that they did not form from the same pulse of fluid/melt as the innermost zone. In
471 addition, N abundance within the rim of diamond 133-6 is consistently high and systematic
472 variations in stable isotope composition are absent, implying that the final growth stage did not
473 occur in a fluid/melt-limited system, in contrast with the innermost growth zone showing
474 systematic N and δ¹³C variations that point to a fluid/melt-limited system.

475 Diamond 138-7 has overall high N concentrations, except for minor growth zones that have
476 N < 100 at.ppm. The absence of a systematic covariations in δ¹³C and N abundance again suggests
477 diamond formation in a system that was not fluid/melt-limited. Although mild resorption is
478 observed at the outer boundary of the inner growth zone, this diamond has homogeneous CL
479 brightness and shows overall small variations in isotope compositions, suggesting that it either
480 formed from two pulses of a related fluid or during protracted growth from fluids with similar
481 isotope compositions and N contents.

482 **6.4 Sources of C and N in Koidu diamonds**

483 The variation of δ¹³C–δ¹⁵N in Koidu diamonds indicates the presence of three major
484 clusters (Fig. 9). Cluster 1 is characterised by depletion in ¹³C and highly variable δ¹⁵N values
485 (δ¹³C = -33.2 to -14.4 ‰; δ¹⁵N = -5.3 to +10.1 ‰); it comprises diamonds with eclogitic inclusions
486 and mixed paragenesis diamond 130-9 (Fig. 13a). Cluster 2 shows mantle-like to mildly ¹³C-
487 enriched C isotope compositions and again a large spread of δ¹⁵N values (δ¹³C = -6.0 to -1.1 ‰;
488 δ¹⁵N = -4.2 to +9.7 ‰); it includes mainly peridotitic diamonds and three mixed paragenesis

489 diamonds – 133-6 (core zone only), 130-31 and 146-1 (Fig. 13a). Cluster 3 has a mantle-like
490 isotopic signature ($\delta^{13}\text{C} = -7.8$ to -3.6 ‰; $\delta^{15}\text{N} = -7.9$ to -2.1 ‰) and includes the rims of eclogitic
491 diamonds and mixed paragenesis diamonds 138-7 and 133-6 (rim only) (Fig. 13b). The assignment
492 of mixed paragenesis diamonds to either Cluster 2 or 3 is entirely based on their N isotope
493 composition, with their $\delta^{13}\text{C}$ values remaining approximately constant over the entire $\delta^{15}\text{N}$ range
494 from ~ -10 to $+10$ ‰.

495 *Cluster 1:* The ^{13}C -depleted signature of our Koidu eclogitic diamonds is also observed in
496 previously studied sulphide-bearing eclogitic diamonds from Koidu ($\delta^{13}\text{C} = -33.0$ to -16.3 ‰;
497 Deines and Harris, 1995). In a suite of five sulphide-bearing eclogitic diamonds from Zimmi
498 (Sierra Leone), which are alluvial diamonds thought to derive from a subcontinental lithospheric
499 mantle source similar to Koidu, three show similar ^{13}C -depleted signatures ($\delta^{13}\text{C} = -24.5$ to -16.1
500 ‰) while the remaining two have $\delta^{13}\text{C}$ between -8.3 and -6.7 ‰ (Smit et al., 2019b). Since $\delta^{15}\text{N}$
501 was not analysed in these two previous studies, it is not possible to assign these diamonds to the
502 three clusters defined in this study. The much wider range of $\delta^{13}\text{C}$ values of eclogitic diamonds
503 compared to peridotitic diamonds has been attributed to fractionation of mantle C during CO_2 -
504 escape from a carbonated fluid/melt before diamond formation (Cartigny et al., 2001). This
505 fractionation model, however, is not likely to produce diamonds with $\delta^{13}\text{C} < -14$ ‰ from a source
506 with initial mantle-like C isotope compositions (Smart et al., 2011). A more probable source of
507 ^{13}C -depleted C for Koidu eclogitic diamonds is recycled crustal material (Milledge et al., 1983;
508 Kirkley et al., 1991), in particular biogenic carbonate \pm organic matter from subducted altered
509 oceanic crust (AOC; Li et al., 2019).

510 The major reservoirs for crustal C in AOC include normal marine carbonate ($\delta^{13}\text{C} \sim 0$ ‰;
511 Schidlowski, 2001; Cartigny et al., 2014), ^{13}C -depleted biogenic carbonate ($\delta^{13}\text{C}$ as low as -24 ‰;

512 Li et al., 2019) and organic matter (average $\delta^{13}\text{C}$ of -26 ± 7 ‰; Schidlowski, 2001). Similarly,
513 crustal N from clay minerals formed by low-temperature ($< 100^\circ\text{C}$) and high-temperature ($> 250^\circ\text{C}$)
514 alteration of oceanic crust has high $\delta^{15}\text{N}$ and low $\delta^{15}\text{N}$ values, respectively (Busigny et al., 2005;
515 Li et al., 2019). Devolatilization of C and N from AOC during subduction further decreases $\delta^{13}\text{C}$
516 and increases $\delta^{15}\text{N}$ values in the residue (Bebout and Fogel, 1992; Li et al., 2014; Cartigny et al.,
517 2014). Devolatilization and mixing of C and N from these various AOC reservoirs (\pm a mantle
518 component) can explain the diverse $\delta^{13}\text{C}$ and $\delta^{15}\text{N}$ values of the eclogitic diamonds contained in
519 Cluster 1 (Fig. 13a).

520 *Cluster 3:* Mantle-derived fluids/melts (i.e., fluids/melts that ultimately derive from the
521 convecting upper mantle) associated with kimberlitic or carbonatitic magmatism are invoked for
522 the crystallization of fibrous diamonds (Boyd et al., 1987; Navon et al., 1988; Boyd et al., 1994b).
523 The confined ranges of $\delta^{13}\text{C}$ and $\delta^{15}\text{N}$ values (both are close to -5 ‰; Boyd et al., 1987; Cartigny,
524 2005; Cartigny et al., 2014; Petts et al., 2016) in fibrous diamonds worldwide indicate that the
525 mantle reservoir of their parental fluids/melts has homogeneous C and N isotope compositions.
526 The $\delta^{13}\text{C}$ and $\delta^{15}\text{N}$ values of rocks that are sourced directly from the convecting upper mantle, i.e.
527 fresh mid-ocean ridge basalts (MORB), also fall in similar ranges (Marty and Zimmermann, 1999;
528 Cartigny et al., 2014), suggesting that they are the isotopic signature of the convecting depleted
529 mantle. Thus, Koidu diamonds in Cluster 3, with C and N isotope compositions similar to fibrous
530 diamonds and MORB, likely formed from relatively homogeneous, mantle-derived fluids/melts
531 (Fig. 13b).

532 The observation that the eclogitic diamond rims fall into “mantle-like” Cluster 3 suggests
533 that the diamond-forming melts/fluids switched from slab-derived to mantle-derived. Decoupling
534 between mineral inclusion composition (related to the eclogitic diamond substrates) and diamond

535 C and N (derived externally, from deeper portions of a slab and/or the lithospheric or convecting
536 mantle) is common, as revealed by a prominent mode in $\delta^{13}\text{C}$ at -5 ‰ and mostly negative $\delta^{15}\text{N}$
537 values for eclogitic inclusion-bearing diamonds worldwide (Cartigny, 2005; Stachel et al., 2009;
538 Cartigny et al., 2014). Although mixed paragenesis diamond 138-7 (in Cluster 3) contains both
539 eclogitic and peridotitic inclusions, none of its growth zones show isotope compositions clearly
540 indicative of subducted C and N, again indicating decoupling between an evolving/changing
541 substrate and the apparently constant source of the diamond-forming fluid.

542 *Cluster 2:* Mixing of C and N derived from the convecting mantle with subducted
543 components shifts diamond $\delta^{13}\text{C}$ and $\delta^{15}\text{N}$ away from the mantle value, following mixing arrays
544 that are controlled in their curvature by the ratio $(\text{N/C})_{\text{Mantle}}/(\text{N/C})_{\text{AOC}}$ (Li et al., 2019) (Fig. 13a).
545 Almost two-thirds of the diamonds in Cluster 2 have $\delta^{13}\text{C} > -4$ ‰ and the majority have positive
546 $\delta^{15}\text{N}$ values, possibly reflecting variable degrees of mixing between a mantle-derived component
547 ($\delta^{13}\text{C} = -5$ ‰ and $\delta^{15}\text{N} = -5$ ‰) and a normal marine carbonate- ($\delta^{13}\text{C} \approx 0$ ‰) and low-temperature
548 clay-derived ($\delta^{15}\text{N} > 0$ ‰) subducted component (Li et al., 2019). Generally, diamonds showing
549 only minor ^{13}C enrichment relative to the mantle value but strongly variable $\delta^{15}\text{N}$ values suggest
550 decoupling of C and N (Mikhail et al., 2014). Alternatively, the slightly ^{13}C -enriched character of
551 Cluster 2 could be a mantle signature unique to the West African Craton, since similar mild shifts
552 to elevated $\delta^{13}\text{C}$ values were also observed for peridotitic diamonds from Kankan (Guinea; Stachel
553 et al., 2002) and Akwatia (Ghana; Stachel and Harris, 1997). The observation that this signature
554 extends to lower mantle diamonds from Kankan (Stachel et al., 2002; Palot et al., 2014) may
555 indicate a plume source of mildly ^{13}C -enriched C, possibly ultimately linked to deeply subducted
556 crustal material. The spread in N isotope composition of Cluster 2 could then also be linked to

557 mixing between asthenosphere-derived ($\delta^{15}\text{N}$ of -5 ± 2 ‰; Cartigny and Marty, 2013) and plume-
558 derived components ($\delta^{15}\text{N}$ of $+3 \pm 2$ ‰; Dauphas and Marty, 1999; Marty and Dauphas, 2003).

559 **6.5 Growth episodes of Koidu diamonds**

560 The three distinct clusters in $\delta^{13}\text{C}$ – $\delta^{15}\text{N}$ space suggest that multiple diamond growth events
561 occurred in the lithospheric mantle beneath Koidu (Fig. 13). Eclogitic diamond cores formed from
562 subducted crustal material (Cluster 1) and rims from mantle-derived fluids/melts (Cluster 3). This
563 sequence indicates eclogitic diamond growth in at least two distinct episodes. One mixed
564 paragenesis diamond (133-6) shows inner growth zones indicative of mixed C and N sources
565 (mantle and subducted; Cluster 2) and a rim that again has a purely mantle-like signature (Cluster
566 3). The abrupt change of isotopic signature between core and rim in these diamonds documents
567 that their growth followed two sequences: Cluster 1 \rightarrow Cluster 3 and Cluster 2 \rightarrow Cluster 3, with
568 the temporal relationship between Clusters 1 and 2 remaining unconstrained.

569 Crystallization of the earlier generation of eclogitic diamonds (Cluster 1) may have been
570 coeval with or subsequent to Archean subduction and eclogite emplacement beneath the Man
571 Shield (Barth et al., 2002a; Aulbach et al., 2019a). For eclogitic sulphide-bearing diamonds from
572 nearby Zimmi, Re-Os dating indicated formation only in the Neoproterozoic (~650 Ma; Smit et
573 al., 2016), but involving S that had experienced mass independent isotopic fractionation in the
574 Archean atmosphere (Smit et al., 2019a), thereby documenting an indirect link to Archean
575 subduction processes. Given the subduction signature of Cluster 1 diamonds, their formation likely
576 is associated with C contained within the Archean slab, but an additional fluid/melt pulse may still
577 be required to facilitate mobilization and re-precipitation of the subducted C and to add a mantle-
578 like N component to the compositional array.

579 A detailed study of Koidu eclogites (Aulbach et al., 2019b) indicated that a subset of
580 diamondiferous low-MgO eclogites and gabbroic eclogites were transformed to barren high-MgO
581 eclogites and pyroxenites as a consequence of metasomatic overprint associated with the
582 Neoproterozoic break-up of Rodinia. The 650 Ma age obtained for Zimmi diamonds (Smit et al.,
583 2016) correlates with the timing of this event. Mixed paragenesis diamonds with a mantle-like
584 stable isotope composition, either throughout (138-7) or in a distinct rim zone (133-6), and the
585 mantle-like composition of the secondary overgrowth on eclogitic diamonds suggest that these
586 metasomatic processes and the transition to diamond formation with mantle-like C and N isotopic
587 signatures may be linked.

588 We propose that after initial growth of eclogitic diamonds (Cluster 1), there were at least
589 two pulses of metasomatic agents involved in the growth of diamonds, the first resulting in Cluster
590 2 and the second in Cluster 3. The first pulse of carbonated ultrabasic melt, derived possibly from
591 either a mantle plume or melting initiated by subduction-related fluids, infiltrated eclogite lenses
592 in the lithospheric mantle, converting some low-MgO and gabbroic eclogites to high-MgO
593 eclogites and pyroxenites (Hills and Haggerty, 1989; Barth et al., 2002b). Some of the original
594 diamond content of the low-MgO and gabbroic eclogites may have been destroyed by the
595 infiltrating melt (Aulbach et al., 2019b). The same pulse of carbonated melt also infiltrated
596 surrounding peridotites (possibly evolving into an aqueous CHO fluid upon equilibration with
597 subsolidus harzburgites). This metasomatic event was associated with the precipitation of Cluster
598 2 diamonds (peridotitic suite and the core of mixed paragenesis diamond 133-6) with mild ^{13}C
599 enrichment and variable ^{15}N enrichment. The trend of outward increasing $\delta^{13}\text{C}$ values and
600 decreasing N concentrations in the central portion of diamond 133-6 suggests that this melt pulse

601 was relatively oxidizing (Fig. 11), driven by carbonate reduction, or redox-neutral precipitation
602 from coexisting CH₄ and CO₂ during cooling.

603 A second pulse of melt/fluid with an asthenosphere-derived C and N isotope signature
604 infiltrated the local lithospheric mantle after a significant period of time, documented by the
605 different N aggregation states of cores and rims of some diamonds (Fig. 12). This pulse was more
606 N-rich and a lack of covariation between $\delta^{13}\text{C}$ values and N abundance in precipitated Cluster 3
607 diamonds indicates that fluid-limited conditions did not occur. Infiltration of this melt/fluid is
608 documented in the rims of eclogitic diamonds and some mixed paragenesis diamonds (133-6 (rim)
609 and 138-7).

610 The two instances of peridotitic inclusions in the rims of mixed paragenesis diamonds are
611 both related to this second metasomatic event, based on mantle-like $\delta^{13}\text{C}$ and $\delta^{15}\text{N}$ values and high
612 N contents of these diamond growth zones. As discussed above, the earlier eclogitic and later
613 peridotitic inclusions in the studied mixed paragenesis diamonds have the respective chemical
614 signatures of eclogitic and peridotitic mantle xenolith minerals, requiring that the diamonds
615 involved must have physically moved between different diamond substrates. Infiltration of an
616 eclogite-derived fluid into a peridotitic diamond substrate (Mikhail et al. 2021) is neither consistent
617 with the clearly crust-derived signatures of the eclogitic inclusions nor with the very Mg- and Cr-
618 rich character of the peridotitic inclusions. Physical transport may have been associated with melt
619 injection, as documented by polymict mantle breccias (juxtaposition of minerals and rock clasts
620 from strongly disparate lithologies) observed as xenoliths from the Kaapvaal craton (Zhang et al.,
621 2003; Giuliani et al., 2014). Alternatively, intense deformation involving small eclogite bodies
622 may have achieved transport of resistant minerals such as diamond into surrounding peridotite
623 along high-strain shear zones. Partial melting of eclogites during fluid influx (Spetsius, 1998) and

624 segregation of such melts may also have allowed for physical transport of originally eclogitic
625 diamonds into peridotitic substrates, followed by renewed diamond growth and encapsulation of
626 peridotitic minerals (Wang, 1998). As indicated by a mild resorption boundary within the one
627 mixed paragenesis diamond entirely associated with Cluster 3 (138-7), this second fluid/melt
628 infiltration may have occurred as repeated pulses over an extended period of time.

629

630 **7. Conclusions**

631 Combined C and N isotope composition of diamond is a robust tracer of the source of
632 diamond-forming fluids/melts. The distinct isotopic signatures, CL responses and N aggregation
633 states observed in the cores and rims of Koidu diamonds suggest episodic diamond growth during
634 multiple pulses of melts/fluids. In their stable isotope compositions, Koidu diamonds of peridotitic,
635 eclogitic and mixed paragenesis form three major clusters: Cluster 1 (cores of eclogitic diamonds)
636 has low $\delta^{13}\text{C}$ and highly variable $\delta^{15}\text{N}$ values ($\delta^{13}\text{C} = -33.2$ to -14.4 ‰; $\delta^{15}\text{N} = -5.3$ to $+10.1$ ‰),
637 suggesting derivation from recycled crustal material (\pm a mantle component). Cluster 2 (peridotitic
638 diamonds and core of mixed paragenesis diamond 133-6) has mantle-like to mildly ^{13}C -enriched
639 C isotope compositions and a wide range of $\delta^{15}\text{N}$ values ($\delta^{13}\text{C} = -6.0$ to -1.1 ‰; $\delta^{15}\text{N} = -4.2$ to $+9.7$
640 ‰), likely reflecting mixing of C and N from subducted and mantle sources. Cluster 3 (rims of
641 eclogitic diamonds and of mixed paragenesis diamond 133-6, and mixed paragenesis diamond
642 138-7) has a homogeneous isotopic signature similar to fibrous diamonds and MORB ($\delta^{13}\text{C} = -7.8$
643 to -3.6 ‰; $\delta^{15}\text{N} = -7.9$ to -2.1 ‰), pointing to precipitation from mantle-derived fluids/melts.
644 Transects across mixed paragenesis diamond 133-6 indicate that the fluid/melt pulse responsible
645 for the growth of its innermost zones (Cluster 2) was relatively oxidizing (gradual outward increase
646 in $\delta^{13}\text{C}$ and decrease in N), whereas diamonds in Clusters 1 and 3 likely precipitated in systems

647 that were not fluid-limited, precluding isotopic fractionation, and thus the redox states of their
648 growth media cannot be constrained.

649 Mixed paragenesis diamonds are associated with all three clusters (1, 2 and 3), but the two
650 examples where we could document a transition from an eclogitic paragenesis core to a peridotitic
651 paragenesis rim either show a transition from mixed source-type Cluster 2 (core) to mantle-like
652 Cluster 3 (rim) or fall entirely into Cluster 3. Based on the major and trace element composition
653 of the mineral inclusions in these mixed paragenesis diamonds, which provide good matches to
654 the equivalent minerals in a subset of Koidu eclogites and in cratonic peridotites, we exclude
655 precipitation during a single intense metasomatic event but instead invoke physical transport of
656 diamonds from eclogitic to peridotitic substrates between separate growth events. The highly
657 depleted character of the peridotitic inclusions precludes that diamond transport involved
658 significant interaction between the peridotitic substrates and either eclogite or melt, which favours
659 either mechanical shearing of small eclogite pods residing in lithospheric peridotites or diamond
660 transport and injection into peridotite through small melt volumes.

661

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669

670 **Appendix A. Supplementary Material**

671 Characteristics of Koidu diamonds are shown in Table S1. Data of FTIR analysis of Koidu
672 diamond 128-7 are shown in Table S2. Major element compositions of mineral inclusions in other
673 Koidu mixed paragenesis diamonds are shown in Table S3. Data of SIMS analysis of Koidu
674 diamonds are shown in Table S4. Examples of common surface textures observed in Koidu
675 diamonds are shown in Fig. S1.

676

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- 929

Table 1. Major element (wt%) compositions of mineral inclusions in diamond 133-6 and 138-7.

Sample	133-6*	133-6-4	138-7-1	138-7-2	138-7-3	138-7-4	138-7-6	138-7-7
Mineral	omphacite	Mg-chromite	garnet	garnet	garnet	garnet	garnet	olivine
Paragenesis	E	P	E	E	E	E	E	P
SiO ₂	55.75	< 0.02	40.97	41.85	41.00	39.46	40.26	41.98
TiO ₂	0.26	0.06	0.45	0.45	0.47	0.47	0.47	< 0.02
Al ₂ O ₃	18.94	7.32	23.21	21.94	22.98	23.86	23.27	< 0.02
Cr ₂ O ₃	0.01	64.46	0.04	0.04	0.02	0.03	0.02	0.07
FeO ^{total}	1.93	12.96	9.60	9.56	8.99	8.95	9.11	5.51
NiO	0.03	0.08	< 0.02	< 0.02	< 0.02	0.01	0.01	0.35
MnO	0.02	0.16	0.17	0.18	0.16	0.17	0.16	0.08
MgO	4.96	13.66	9.28	9.70	8.47	8.07	8.36	53.47
CaO	8.65	< 0.02	16.89	16.70	18.55	19.26	19.28	0.02
Na ₂ O	8.33	< 0.02	0.18	0.17	0.18	0.18	0.18	< 0.02
K ₂ O	0.22	< 0.02	< 0.02	< 0.02	< 0.02	< 0.02	< 0.02	< 0.02
Total	99.10	98.70	100.79	100.56	100.82	100.46	101.12	101.48
Mg#	82.2	65.2	63.2	64.4	62.7	61.7	62.1	94.5
Ca#			45.1	44.2	49.5	51.2	50.5	

Mg# = $100 \times \text{Mg} / (\text{Mg} + \text{Fe})$; Ca# = $100 \times \text{Ca} / (\text{Ca} + \text{Mg} + \text{Fe} + \text{Mn})$.

* Average values of multiple omphacites in diamond 133-6 are reported here.

Table 2. Trace element (ppm) compositions of mineral inclusions in diamond 133-6 and 138-7.

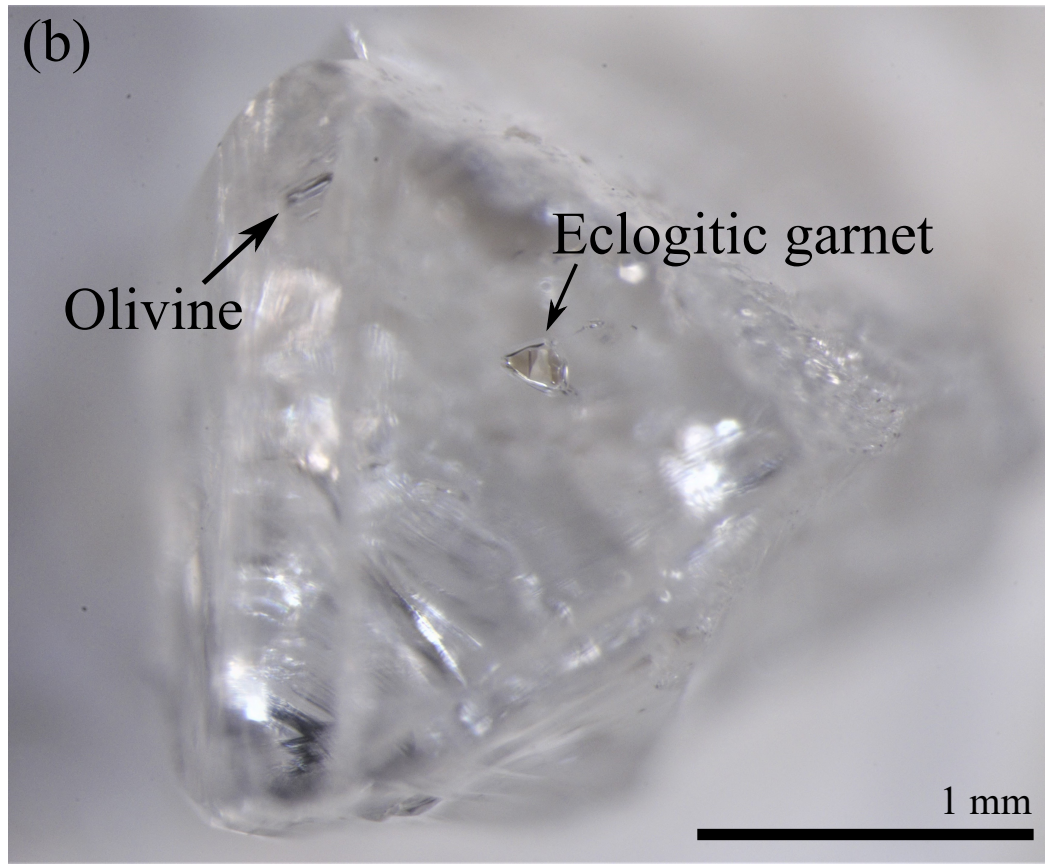
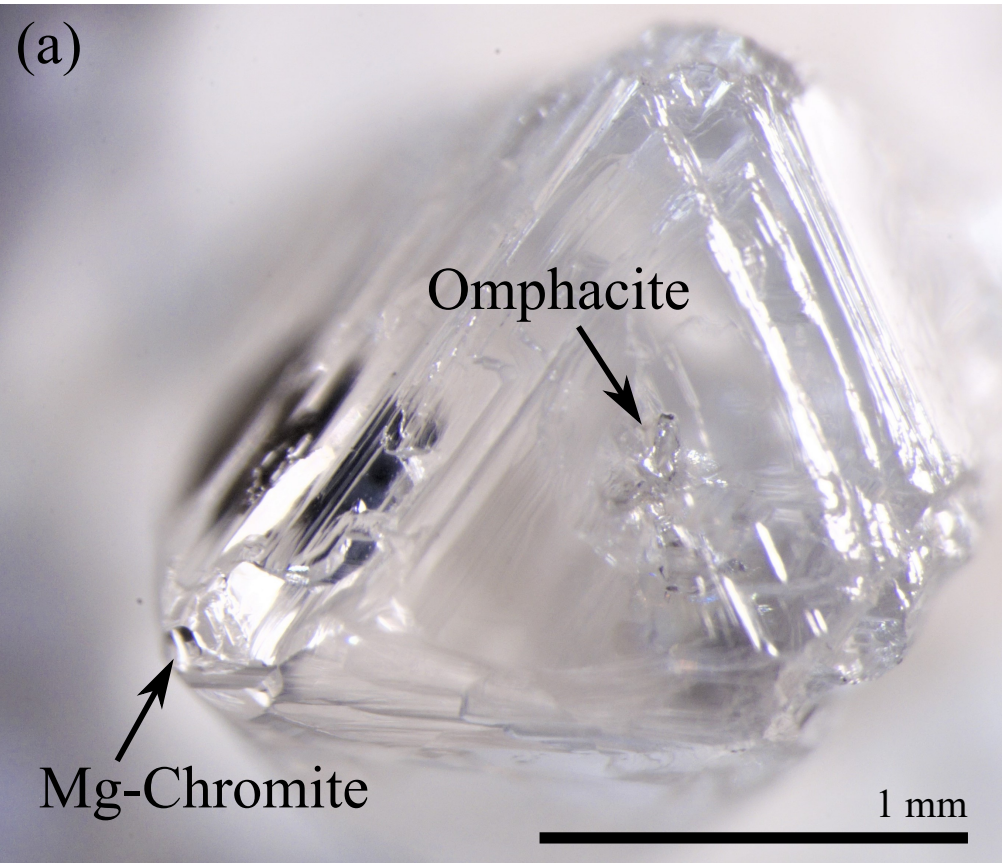
Sample	133-6-1	133-6-2	133-6-3	138-7-1	138-7-3	138-7-6
Mineral	Omphacite	Omphacite	Omphacite	Garnet	Garnet	Garnet
La	2.988	3.000	3.016	0.092	0.062	0.062
Ce	4.480	4.642	4.622	0.644	0.559	0.615
Pr	0.315	0.321	0.317	0.203	0.213	0.215
Nd	0.800	0.827	0.870	2.050	1.910	2.090
Sm	0.154	0.142	0.140	1.300	1.314	1.392
Eu	0.059	0.060	0.069	0.716	0.617	0.854
Gd	0.067	0.088	0.100	1.970	2.240	2.440
Tb	0.005	0.008	0.008	0.373	0.321	0.468
Dy	0.050	0.030	0.027	2.770	2.138	3.120
Ho	0.002	0.006	0.006	0.614	0.540	0.660
Er	0.010	0.009	0.005	1.740	1.646	1.715
Tm	n.d.	n.d.	n.d.	0.252	0.182	0.234
Yb	n.d.	0.005	n.d.	1.470	1.116	1.414
Lu	n.d.	n.d.	n.d.	0.198	0.155	0.201
Sc	1.330	1.560	1.680	41.3	36.1	38.4
Ti	1553	1613	1618	2596	2889	2895
V	108	110	110	86.4	78.5	82.5
Ni	326	321	328	78.3	60.2	65.0
Rb	0.540	0.532	0.510	n.d.	n.d.	n.d.
Sr	135	134	133	5.97	4.02	4.05
Y	0.107	0.116	0.097	15.3	14.5	16.6
Zr	2.2	2.2	2.3	12.2	7.1	12.1
Nb	1.120	1.088	1.106	0.091	0.106	0.142
Ba	0.380	0.685	0.628	n.d.	n.d.	n.d.
Hf	0.137	0.131	0.158	0.233	0.146	0.251

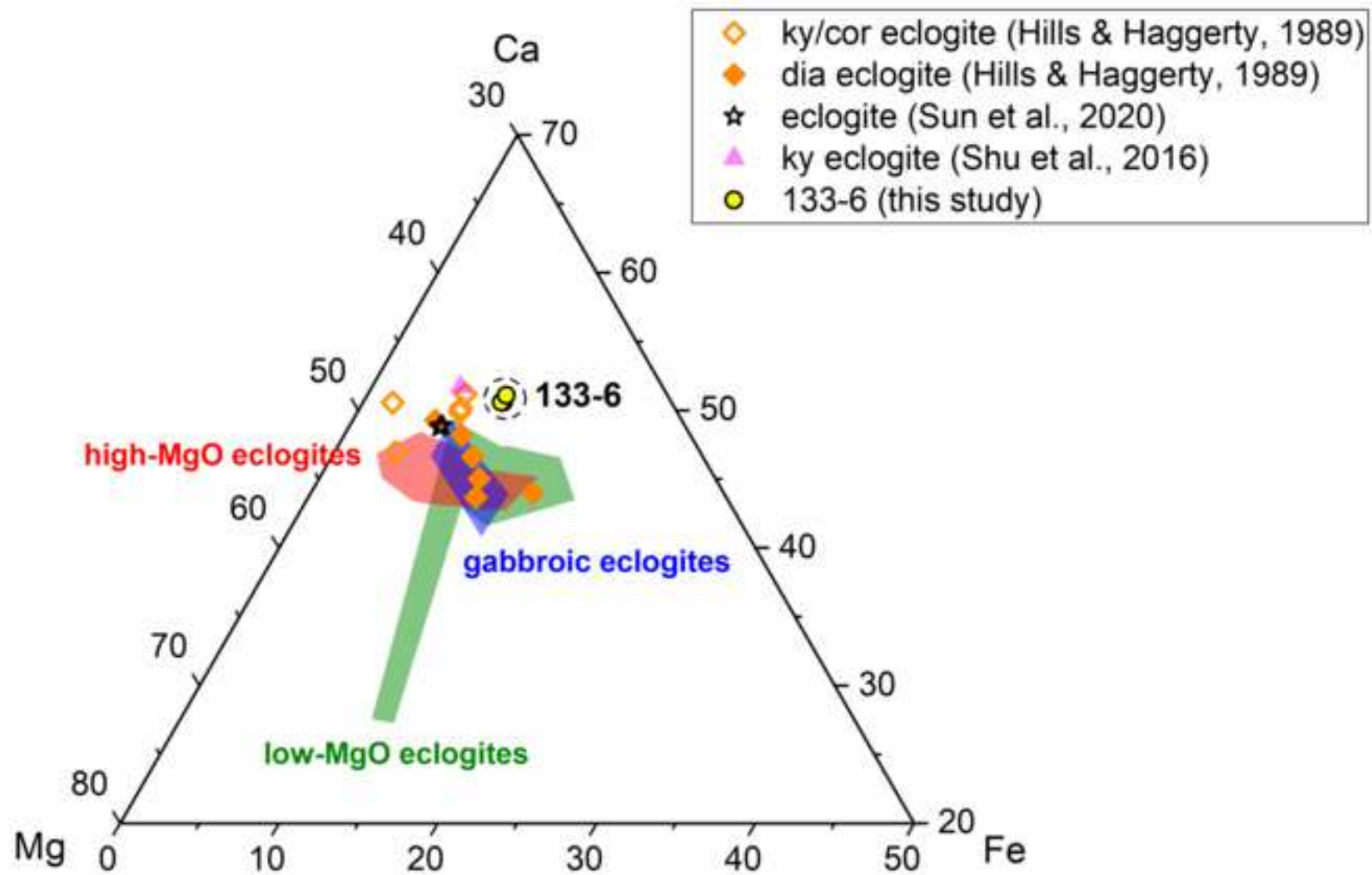
n.d. = not determined.

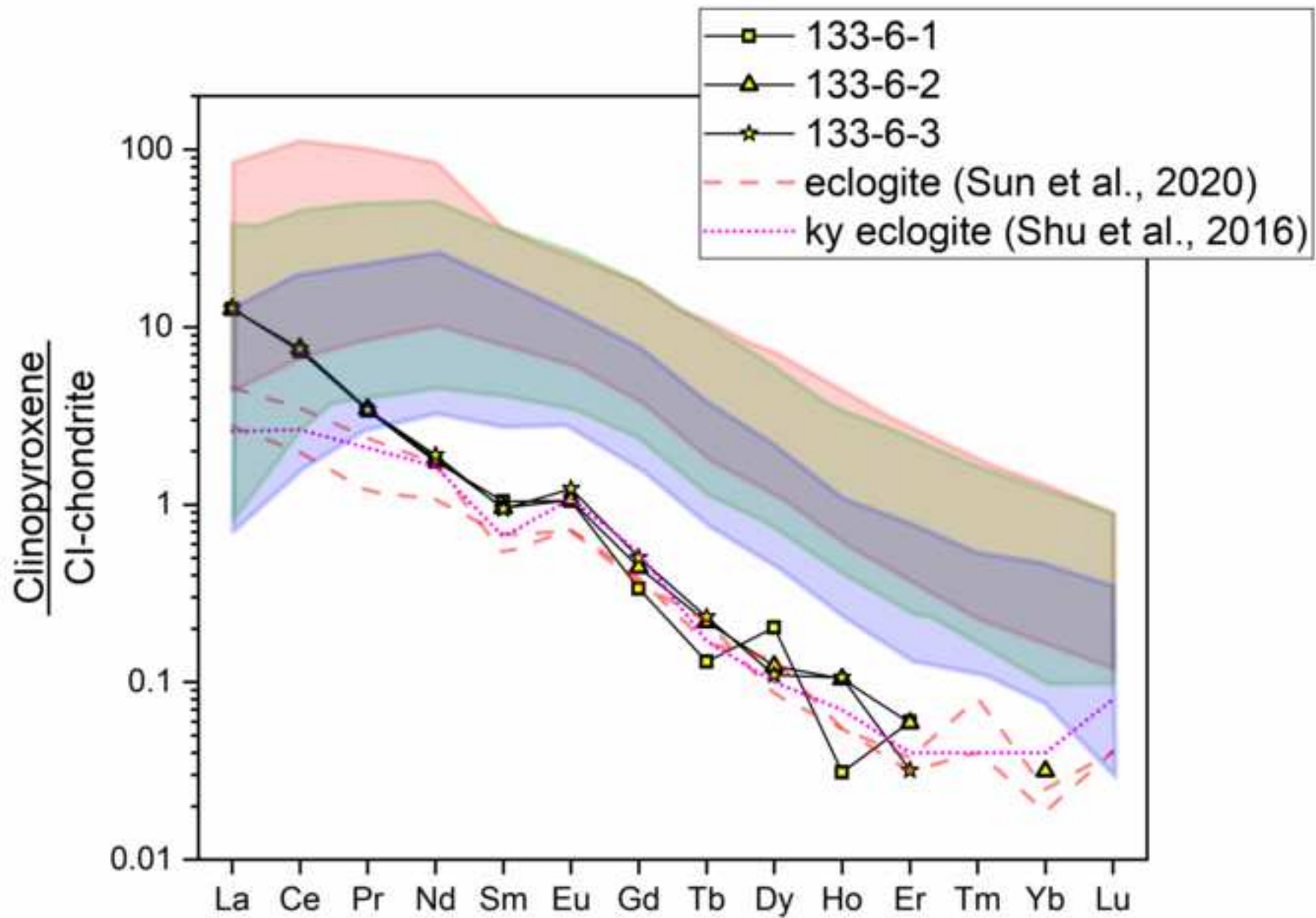
Three omphacites (from diamond 133-6) and three eclogitic garnets (from diamond 138-7) with larger grain size were analyzed.

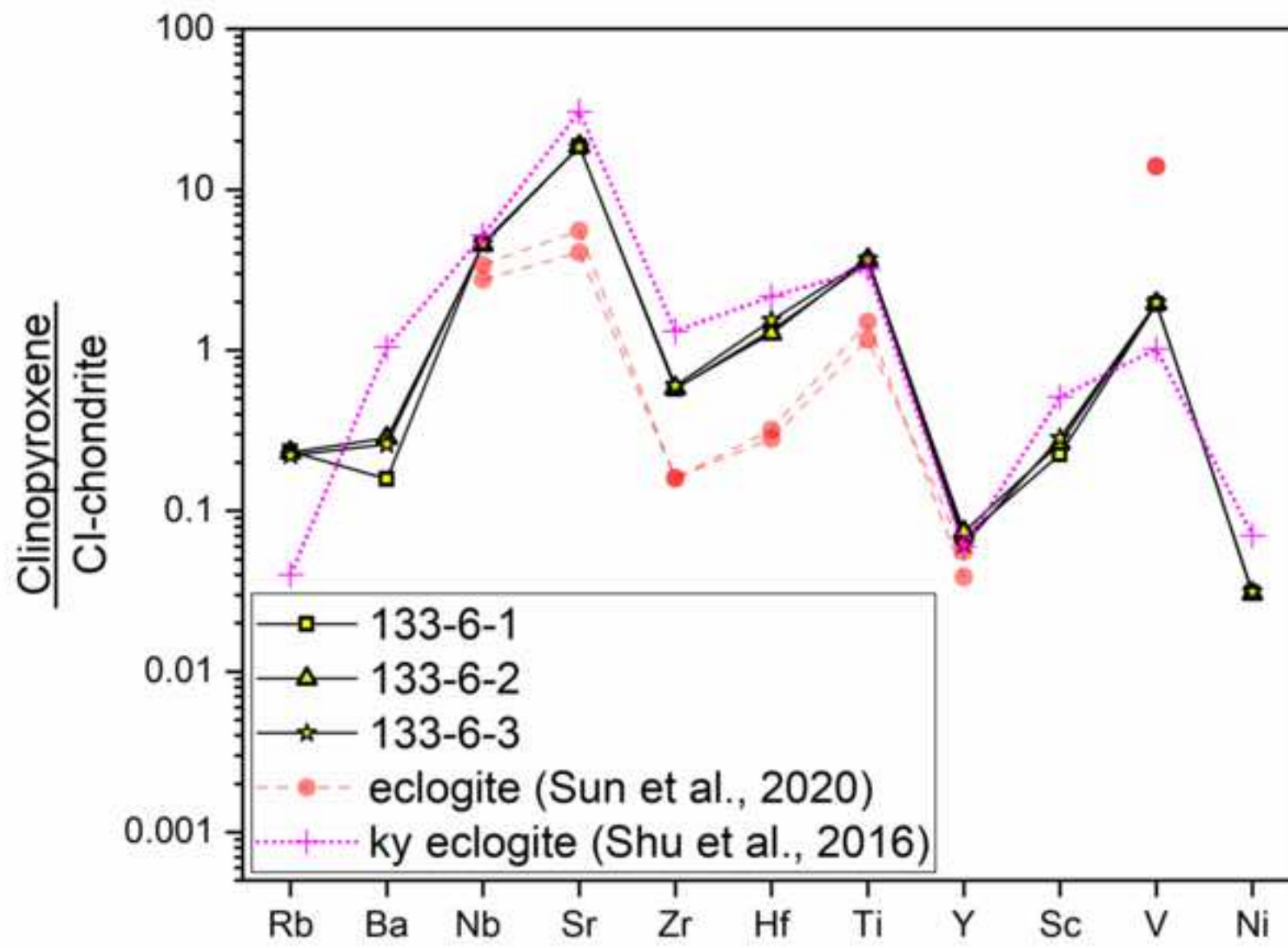
Table 3. Summary of SIMS analysis of Koidu diamonds.

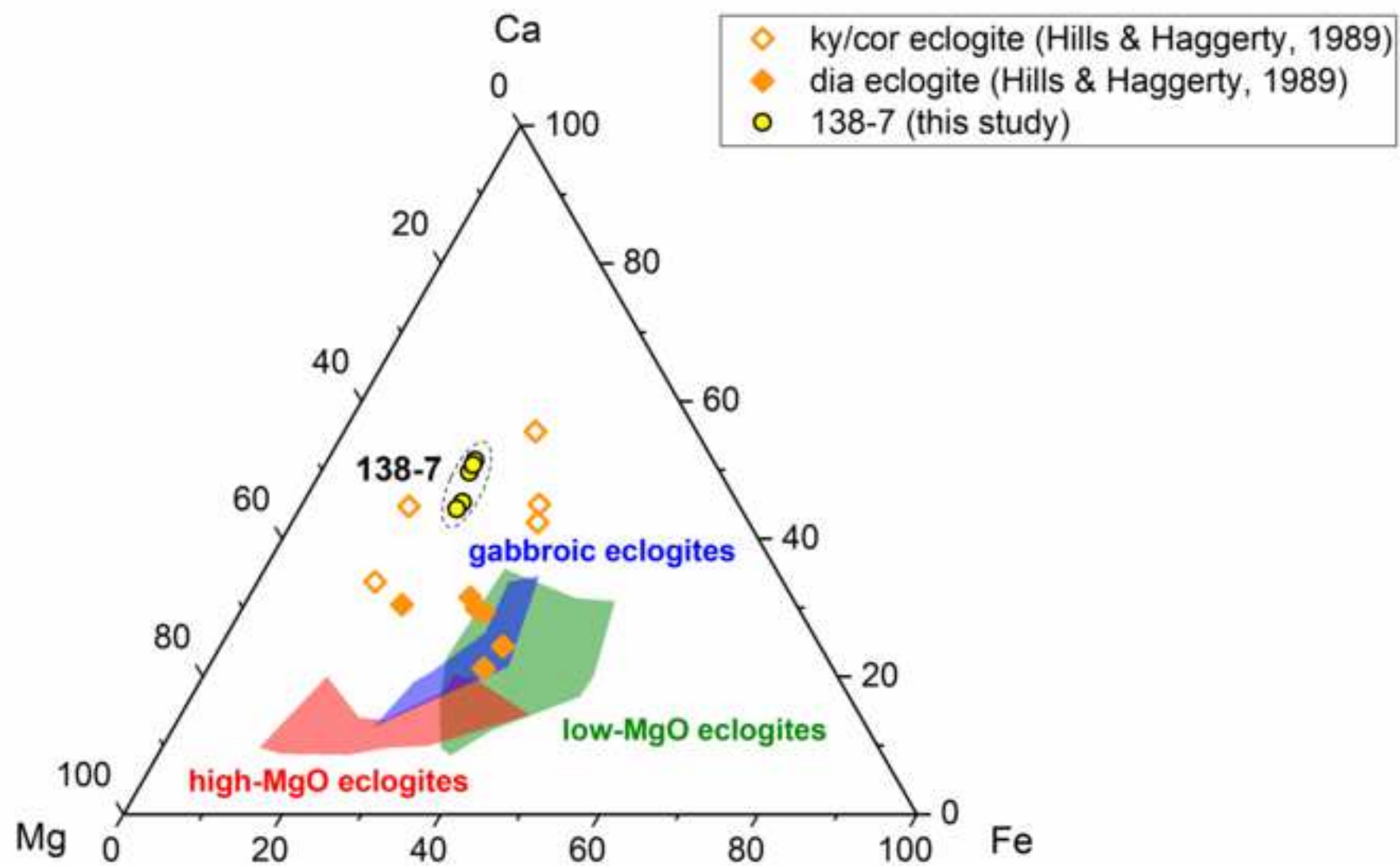
Paragenesis	No. of diamond	$\delta^{13}\text{C}_{\text{VPDB}}$ (‰)	[N] (at.ppm)	$\delta^{15}\text{N}_{\text{Air}}$ (‰)
Peridotitic	18	-6.0 to -1.1	0.4 to 920	-4.2 to +9.7
Eclogitic	82	-33.2 to -0.04	0.4 to 2200	-7.9 to +10.1
Mixed	5	-24.8 to -1.3	0.3 to 1050	-9.8 to +5.8

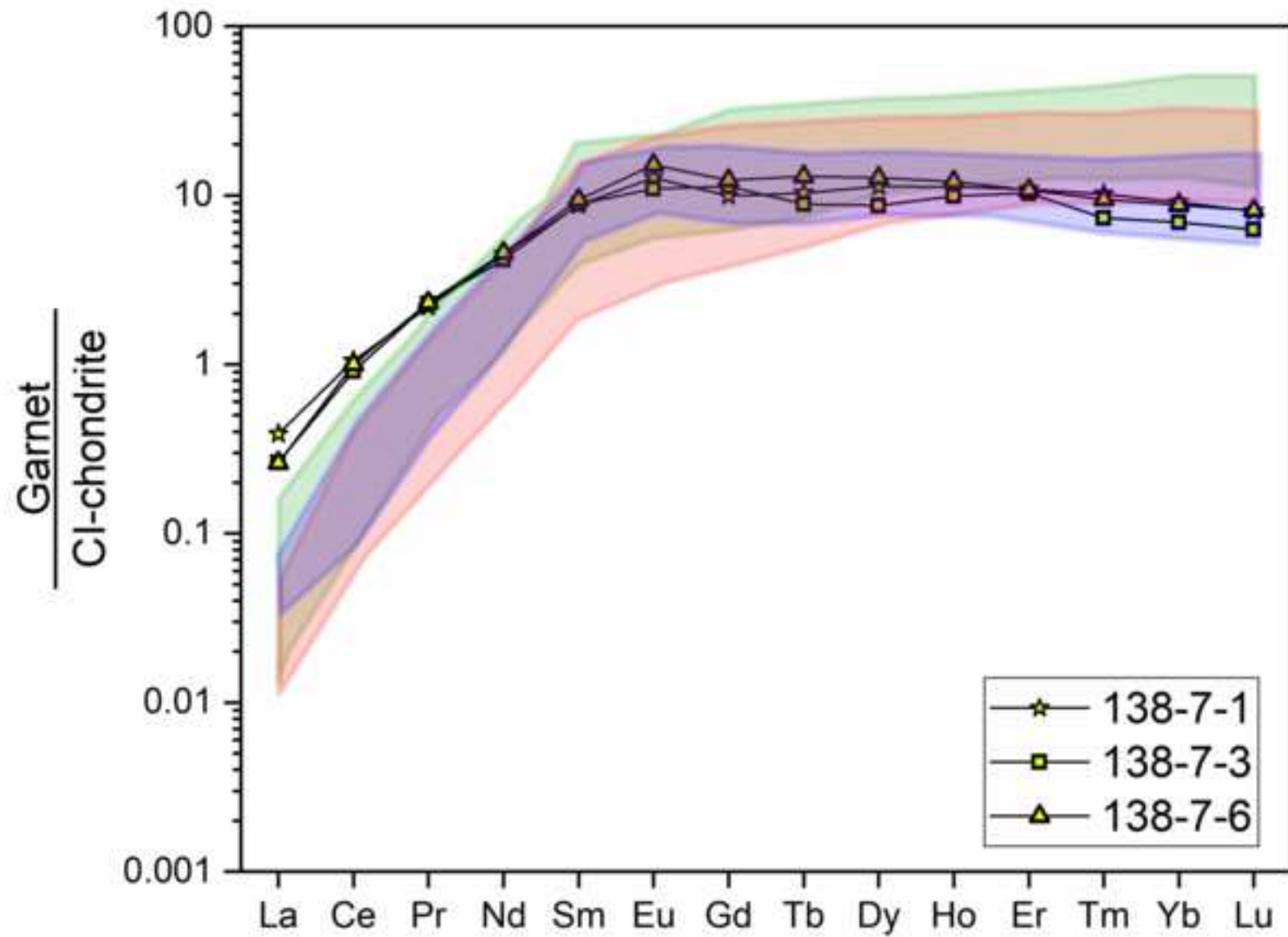


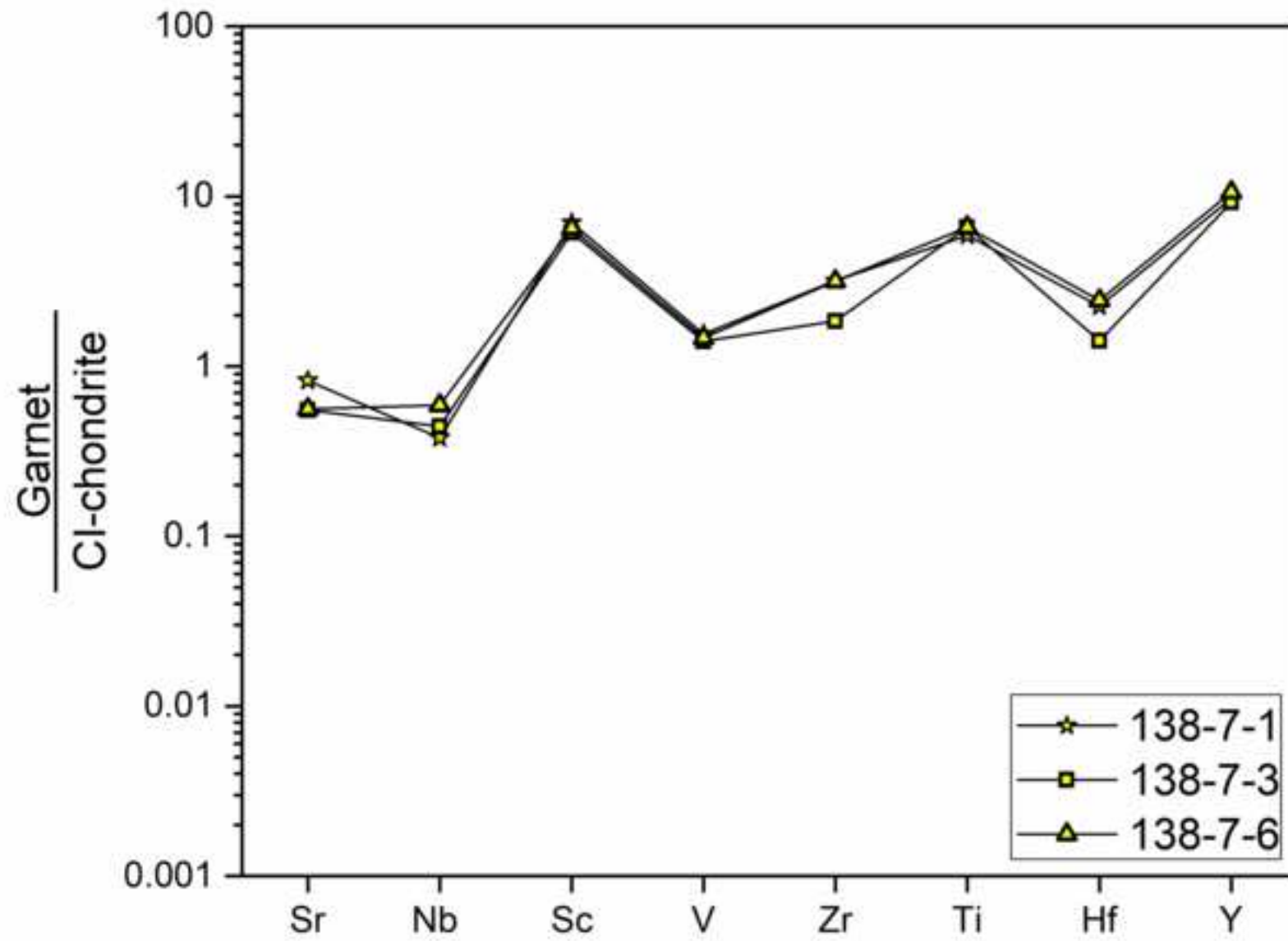












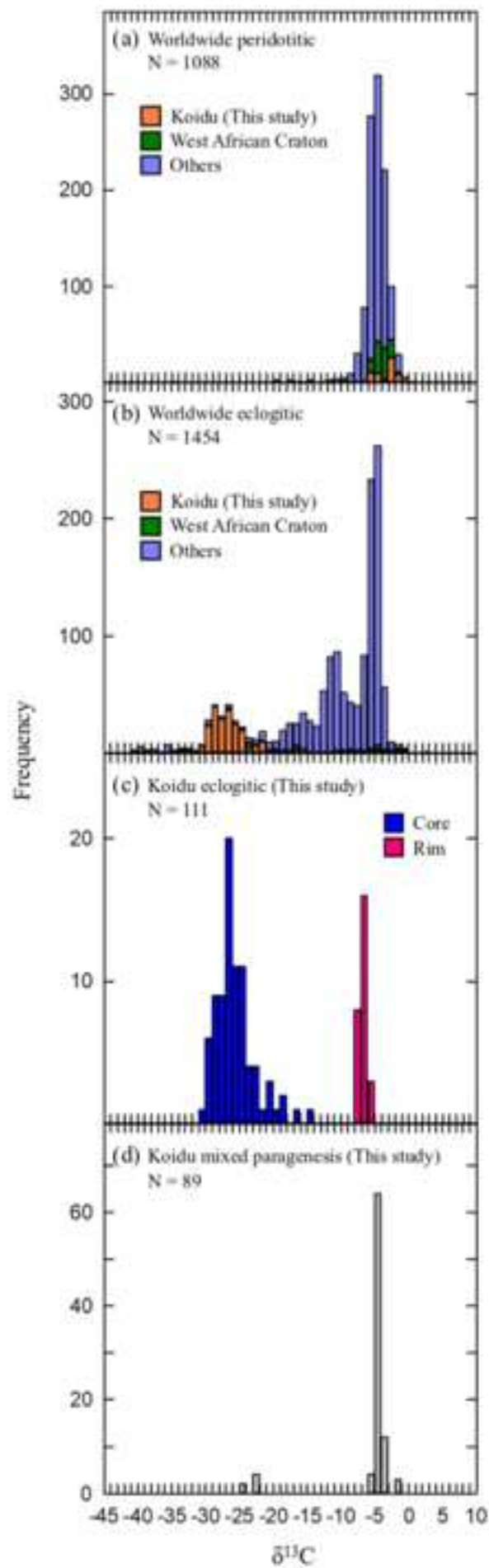
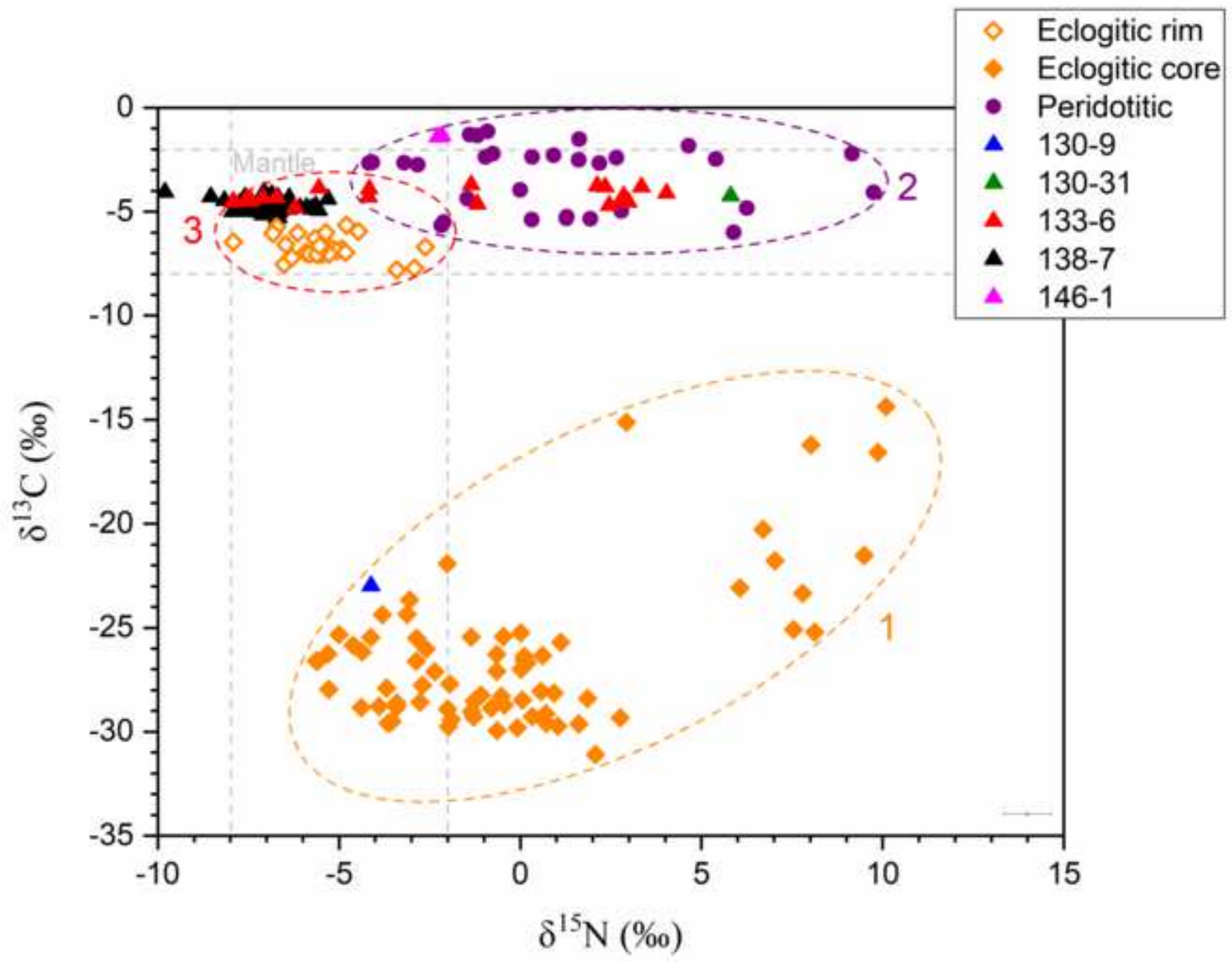
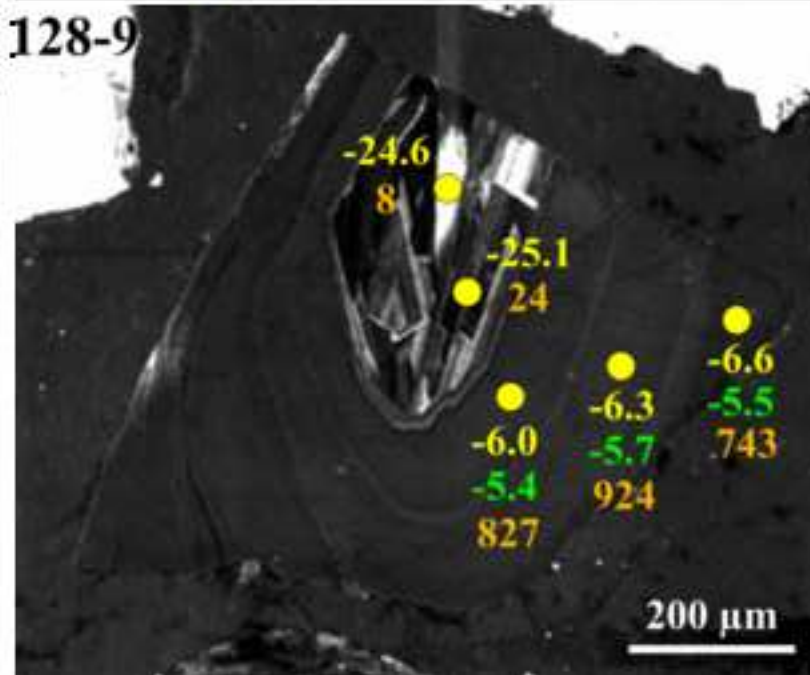
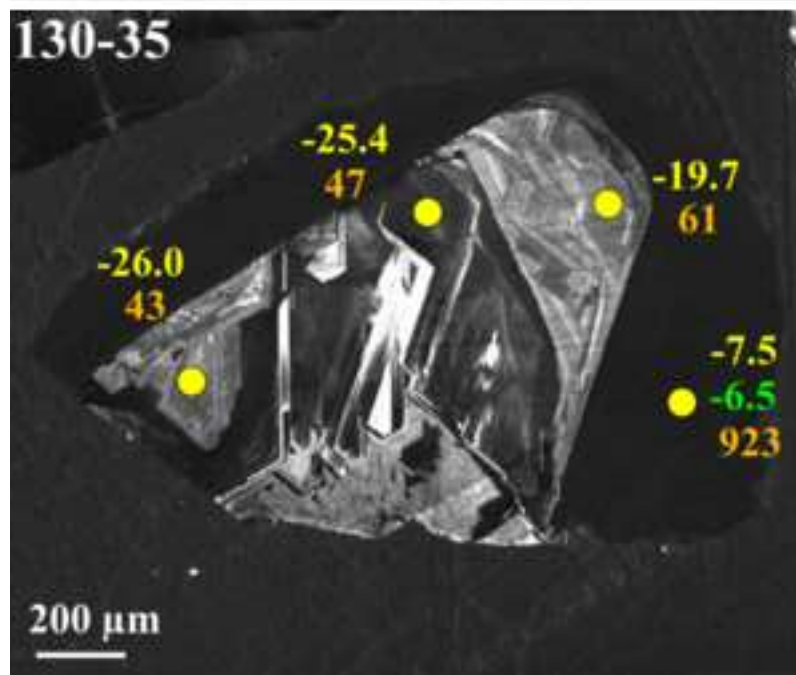
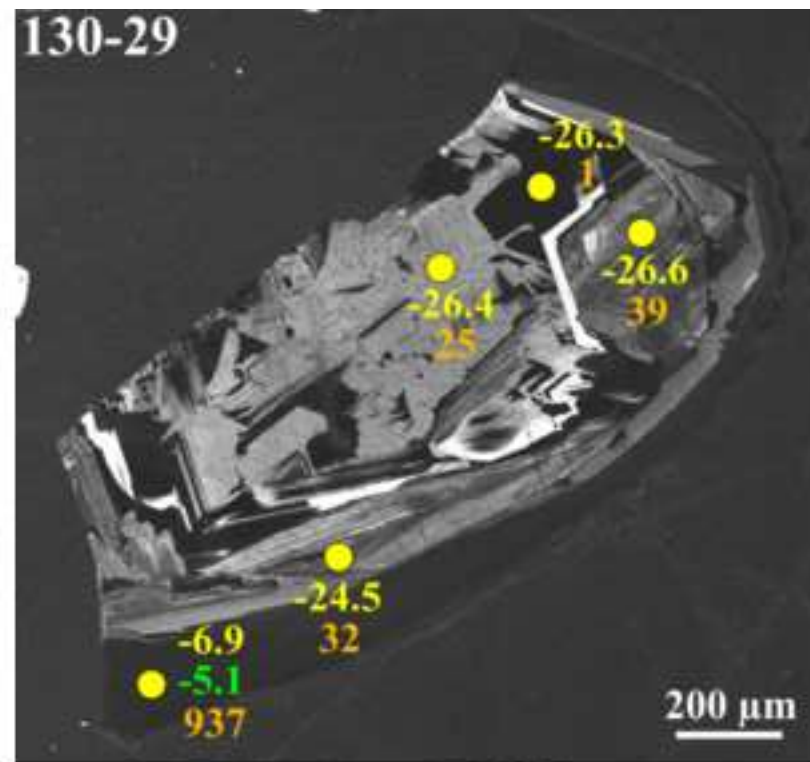
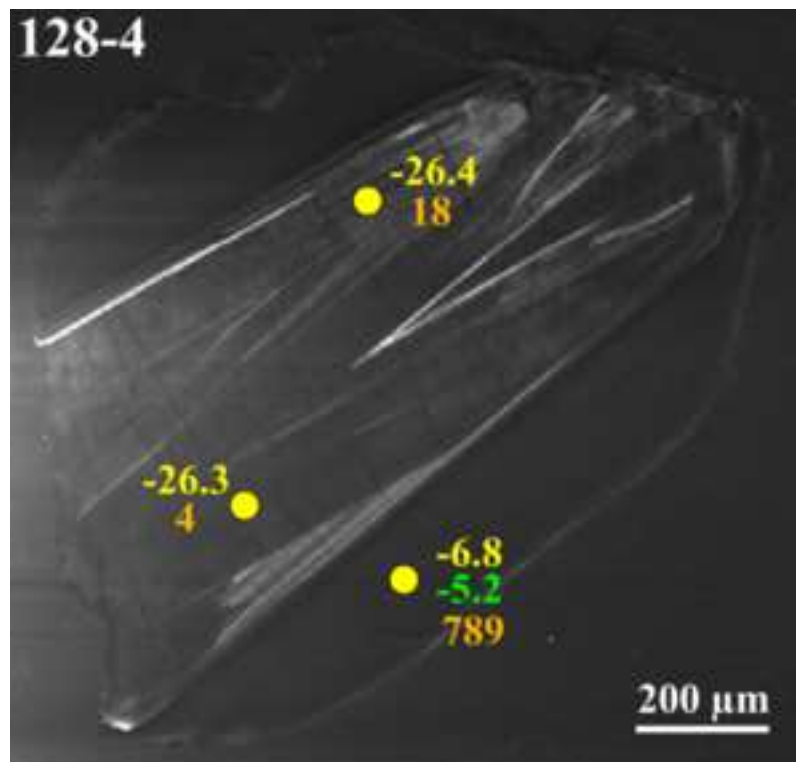
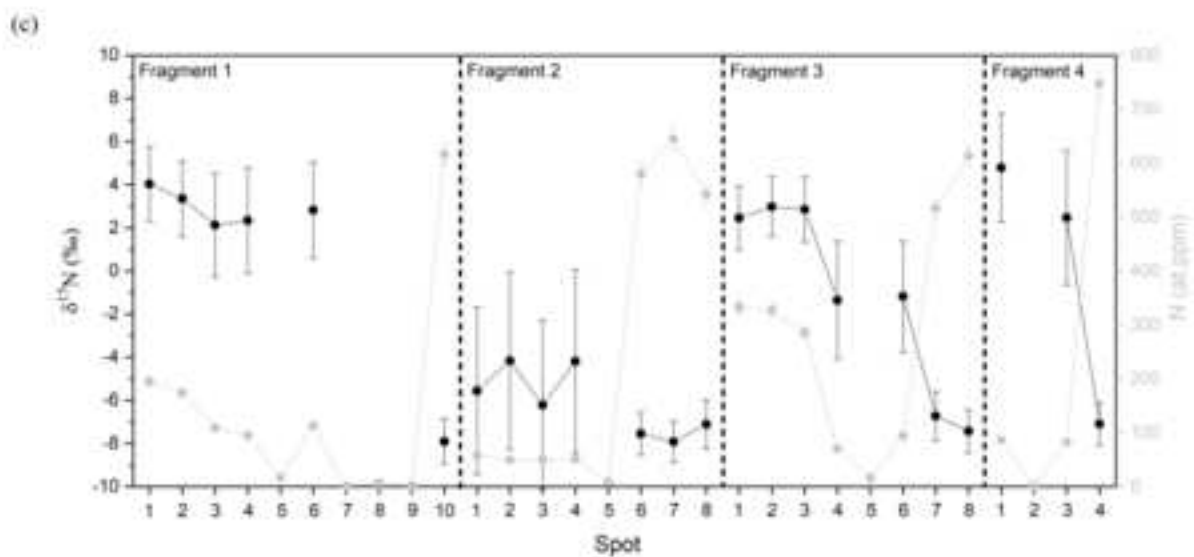
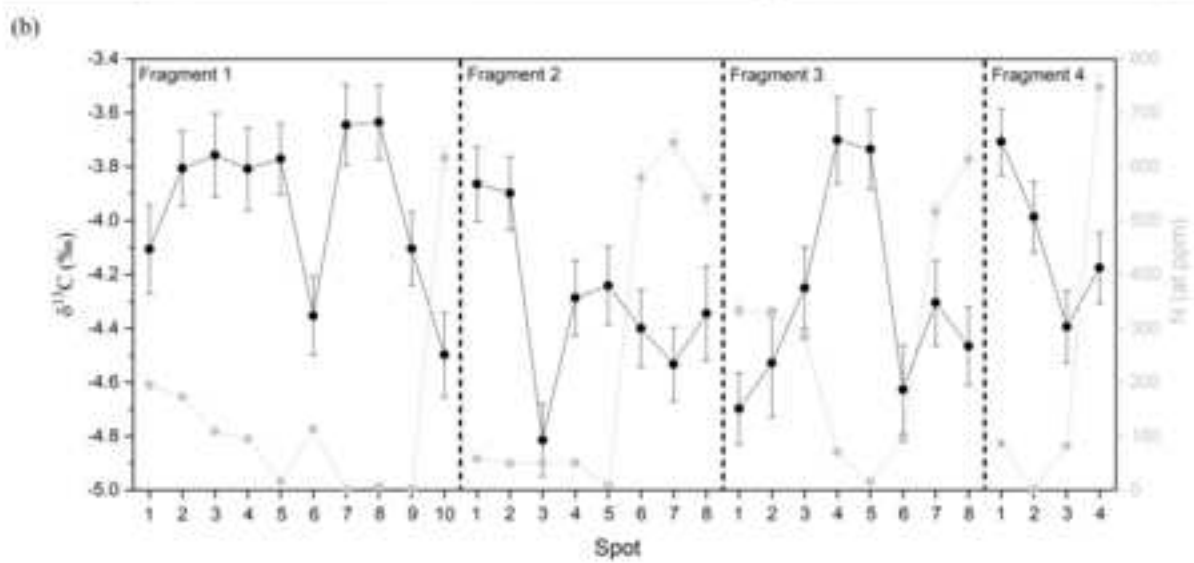
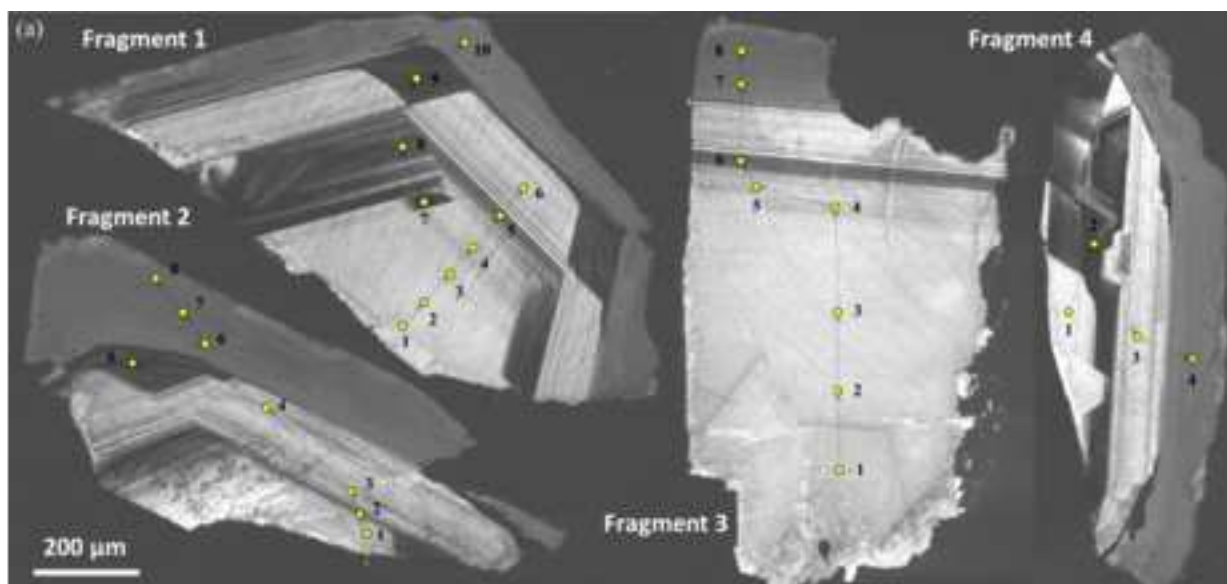
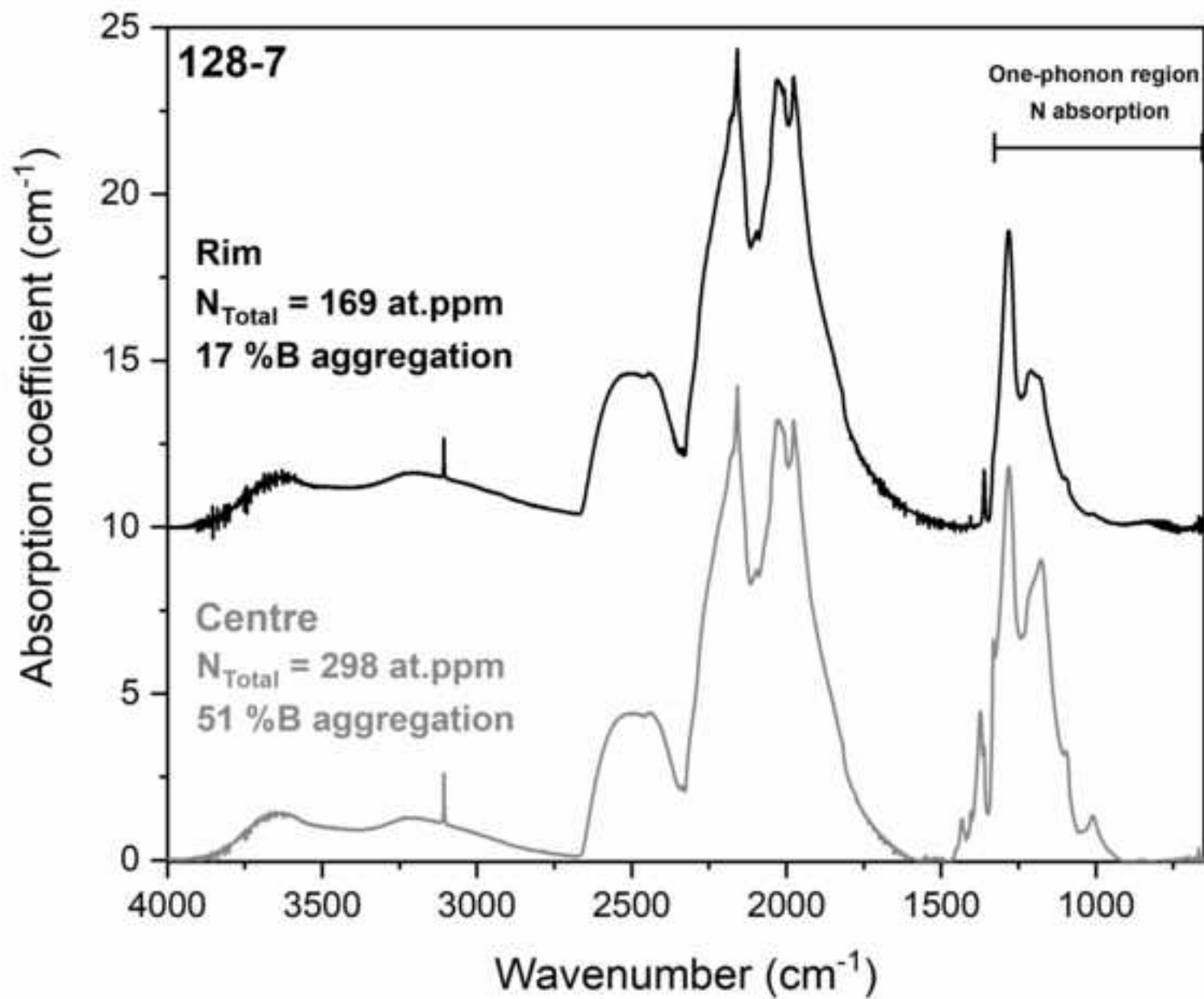


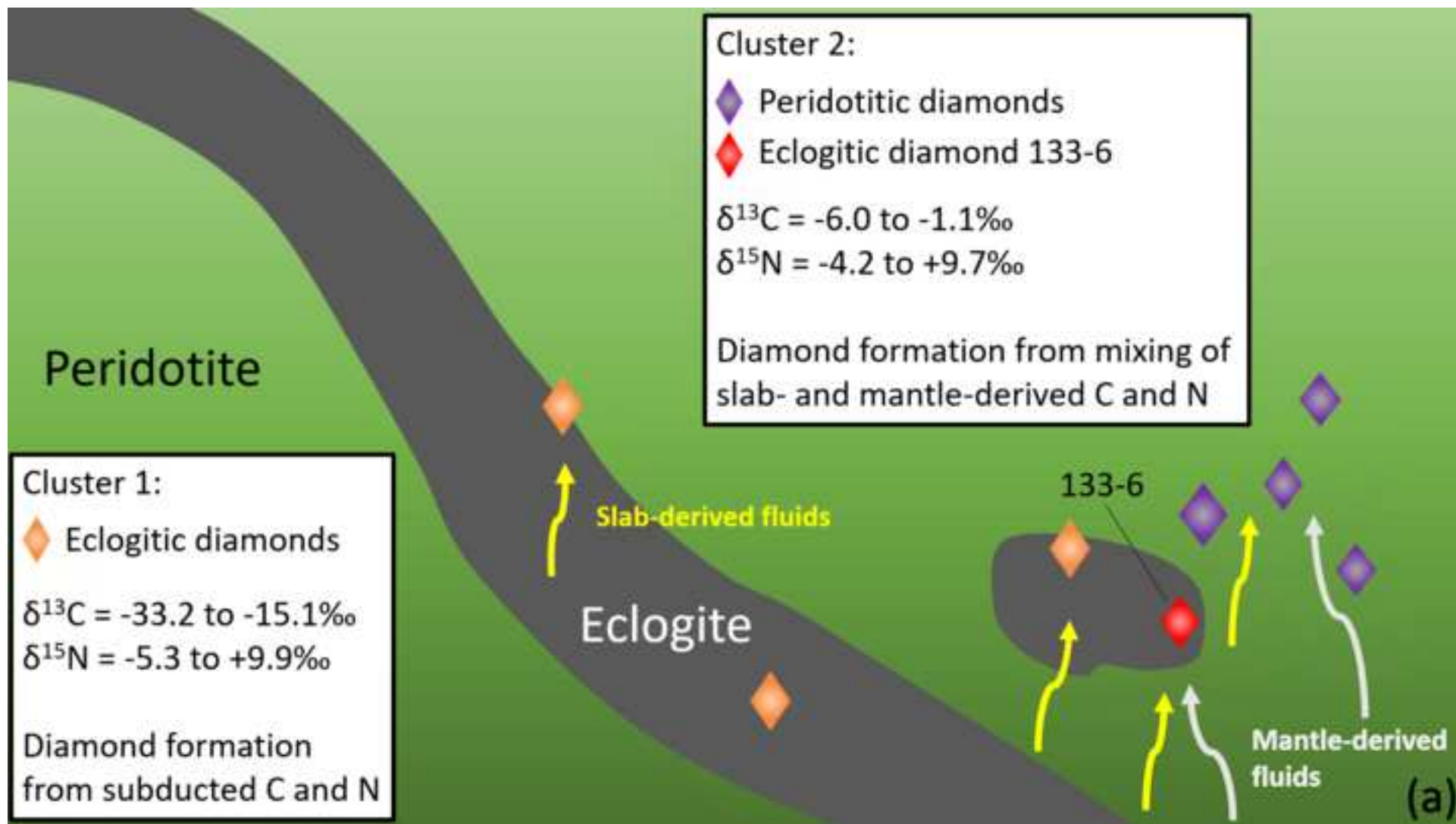
Figure 9











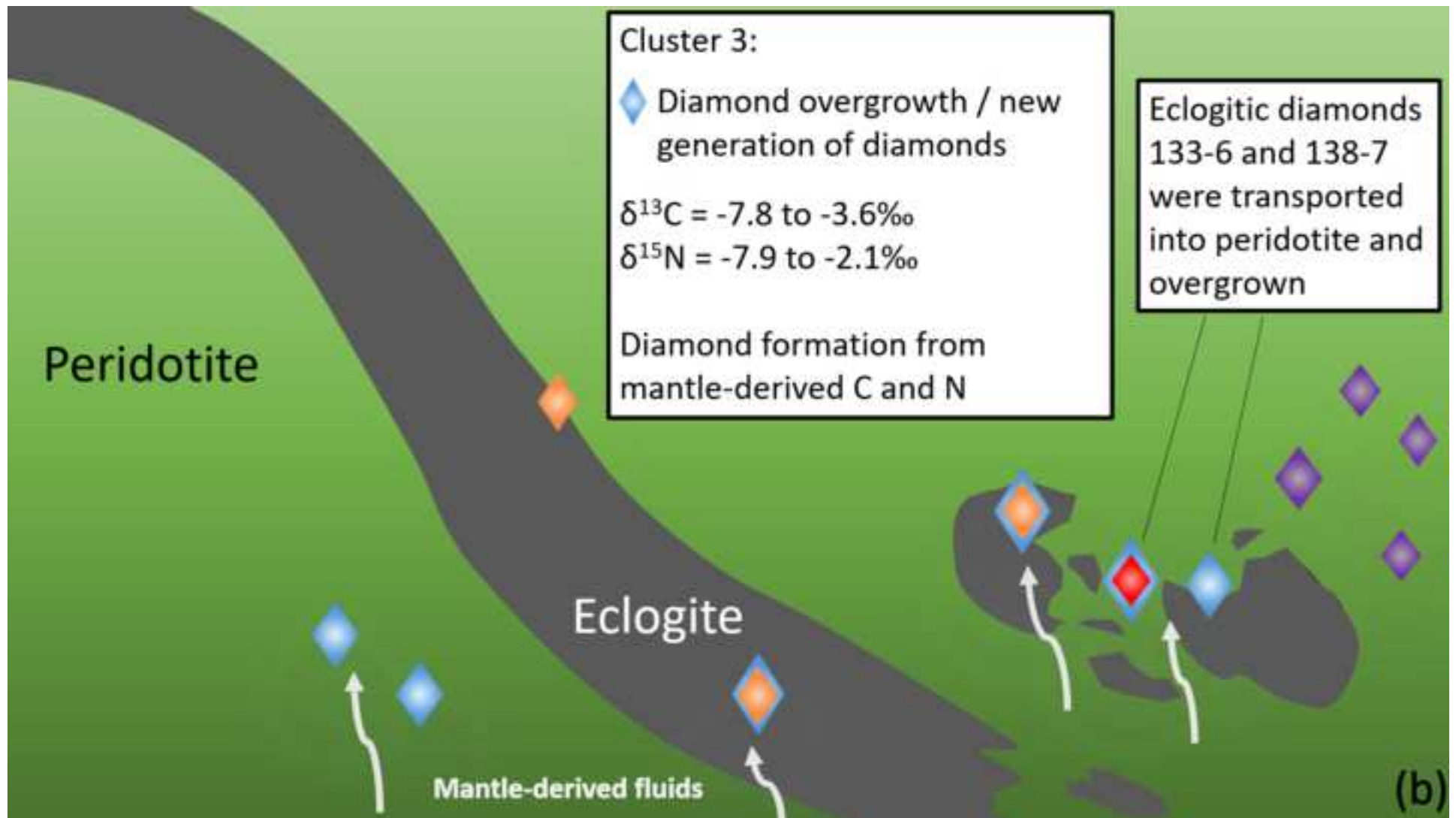


Fig. 1. Diamonds with co-occurring eclogitic and peridotitic mineral inclusions. (a) Diamond 133-6 with multiple omphacites in the centre and a Mg-chromite in the rim. (b) Diamond 138-7 with eclogitic garnets in the centre and an olivine near the rim.

Fig. 2. Ternary diagram showing the major element composition (molar Ca-Mg-Fe) of omphacite inclusions from mixed paragenesis diamond 133-6 in this study (circles). The fields for omphacites from high-MgO (red), low-MgO (green) and gabbroic eclogites (blue) (Aulbach et al., 2019c), and from kyanite- and corundum-bearing low-MgO eclogites (open diamonds) and diamond-bearing low-MgO eclogites (solid diamonds) (Hills and Haggerty, 1989) are shown for comparison. Also shown are omphacites from two low-MgO eclogites from the Obnazhennaya kimberlite in Siberia (overlapping stars; Sun et al., 2020) and from a kyanite eclogite from the Bellsbank kimberlite in South Africa (triangle; Shu et al., 2016).

Fig. 3. REE concentrations in omphacites from mixed paragenesis diamond 133-6, normalized to CI-chondrite (McDonough and Sun, 1995). These omphacites have highly fractionated ($La_N/Yb_N = 422$) REE_N patterns. The fields for omphacites from high-MgO (red), low-MgO (green) and gabbroic eclogites (blue) (Aulbach et al., 2019c) are shown for comparison. Also shown are the REE_N patterns of omphacites from two low-MgO eclogites from the Obnazhennaya kimberlite in Siberia (red dashed lines; Sun et al., 2020) and from a kyanite eclogite from the Bellsbank kimberlite in South Africa (magenta dotted line; Shu et al. 2016), which are very similar to our Koidu samples.

Fig. 4. Concentrations of other trace elements including LILE and HFSE in omphacites from mixed paragenesis diamond 133-6, normalized to CI-chondrite (McDonough and Sun, 1995). Also shown are the omphacites from two low-MgO eclogites from the Obnazhennaya kimberlite in Siberia (red dashed lines; Sun et al., 2020) and from a kyanite eclogite from the Bellsbank kimberlite in South Africa (magenta dotted line; Shu et al., 2016). As already seen for the REE, these omphacites have overall very similar trace element patterns to our samples. Elements are arranged in increasing compatibility in the clinopyroxene structure (Green, 1994).

Fig. 5. Ternary diagram showing the major element compositions (molar Ca-Mg-Fe) of garnet inclusions from mixed paragenesis diamond 138-7 (open circles). The fields for garnets from high-MgO (red), low-MgO (green) and gabbroic eclogites (blue) (Aulbach et al., 2019c) and from kyanite- and corundum-bearing low-MgO eclogites (open diamonds) and diamond-bearing low-MgO eclogites (solid diamonds) (Hills and Haggerty, 1989) are shown for comparison.

Fig. 6. REE concentrations in garnets from mixed paragenesis diamond 138-7, normalized to CI-chondrite (McDonough and Sun, 1995). The fields for garnets from high-MgO (red), low-MgO (green) and gabbroic eclogites (blue) (Aulbach et al., 2019c) are shown for comparison.

Fig. 7. Concentrations of other trace elements including LILE and HFSE in garnets from mixed paragenesis diamond 138-7, normalized to CI-chondrite (McDonough and Sun, 1995). Elements are arranged in increasing compatibility in the garnet structure (Green, 1994).

Fig. 8. Distribution of $\delta^{13}C$ values (individual SIMS analysis spots) in (a) peridotitic diamonds, (b) eclogitic diamonds (only diamonds without a core-rim structure are shown), (c) eclogitic diamonds with a core-rim structure and (d) mixed paragenesis diamonds in this study. Both peridotitic (a) and unzoned

eclogitic (b) Koidu diamonds are compared to diamonds from other localities on the West African Craton and from worldwide sources (database of Stachel et al., 2022).

Fig. 9. $\delta^{13}\text{C}$ versus $\delta^{15}\text{N}$ for peridotitic (circles), eclogitic (solid diamonds: core; open diamonds: rim) and mixed paragenesis diamonds (triangles) from Koidu. The error bars in the lower right corner indicated the typical total analytical uncertainties (95% confidence level). Three major compositional clusters are identified for Koidu diamonds: Cluster 1 ($\delta^{13}\text{C} = -33.2$ to -14.4 ‰; $\delta^{15}\text{N} = -5.3$ to $+10.1$ ‰), Cluster 2 ($\delta^{13}\text{C} = -6.0$ to -1.1 ‰; $\delta^{15}\text{N} = -4.2$ to $+9.7$ ‰) and Cluster 3 ($\delta^{13}\text{C} = -7.8$ to -3.6 ‰; $\delta^{15}\text{N} = -7.9$ to -2.1 ‰).

Fig. 10. Examples of cathodoluminescence images of Koidu eclogitic diamonds (128-4, 128-9, 130-29 and 130-35) showing core–rim structures. SIMS measurement spots with $\delta^{13}\text{C}$ values (‰; yellow), $\delta^{15}\text{N}$ values (‰; green) and N concentrations (at.ppm; orange) are indicated.

Fig. 11. (a) Cathodoluminescence images of four fragments of mixed paragenesis diamond 133-6. The yellow dots indicate the spots of SIMS analyses. (b) Core-to-rim transects across the four fragments showing variations in $\delta^{13}\text{C}$ (‰; black) and N concentration (at.ppm; grey). (c) Core to rim variations across the same fragments for $\delta^{15}\text{N}$ (‰; black) and N concentration (at.ppm; grey). The total uncertainty (95% confidence level) for each point is indicated by error bars unless they are smaller than the size of the symbol.

Fig. 12. Infrared absorption spectra of a Koidu diamond (diamond 128-7) collected through its centre (grey) and rim (black). Nitrogen absorption bands in the one-phonon region (~ 1332 to ~ 400 cm^{-1}) of the spectra show a higher degree of N aggregation (higher percentage of N in B aggregation) in the centre compared to the rim.

Fig. 13. Schematic diagram showing the growth episodes reflected by Koidu diamonds. (a) Cluster 1 diamonds (cores of eclogitic diamonds indicated by orange diamond symbol) formed from subducted crustal material (\pm a mantle component). Cluster 2 diamonds (core of mixed paragenesis diamond 133-6 indicated by red diamond symbol, and peridotitic diamonds indicated by purple diamond symbol) formed from mixing of slab- and mantle-derived C and N. Note that the temporal relationship between Cluster 1 and 2 is unconstrained. (b) Cluster 3 diamonds (rims of eclogitic diamonds and of mixed paragenesis diamond 133-6, mixed paragenesis diamond 138-7 and potentially some peridotitic diamonds indicated by blue diamond symbol) formed from purely mantle-derived C and N. Mixed paragenesis diamonds 133-6 and 138-7 originally formed in eclogite (encapsulating omphacites and eclogitic garnets, respectively) and were then physically transported into surrounding peridotites, where renewed diamond growth encapsulated peridotitic minerals (Mg-chromite and olivine, respectively).