1	Nature of slab-mantle interactions recorded by coupled $\delta^{13}C-$
2	$\delta^{15}N-\delta^{18}O$ signatures and elemental compositions of Koidu
3	diamonds and their inclusions
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Abstract

Eclogitic diamond formation can be associated with interactions between subducted slabs and ambient mantle. To gain a better understanding of the extent of chemical exchange between slabs and ambient mantle and the mechanisms of coeval diamond formation, we examined 16 diamonds with eclogitic garnet inclusions from the Koidu kimberlite complex in Sierra Leone (West African Craton). We analyzed the 16 garnet inclusions for major element, trace element and oxygen isotope (δ^{18} O) compositions and their host diamonds for carbon (δ^{13} C) and nitrogen (δ^{15} N) isotope compositions.

The garnet inclusions have δ^{18} O values ranging from +5.4 to +12.1 ‰ (median = +11.3 ‰) and 26 all but two have δ^{18} O values $\geq +9.9$ ‰. Such high δ^{18} O values indicate a link to protoliths that had 27 undergone extensive low-temperature alteration by seawater, which occurs in the uppermost basaltic 28 layer of oceanic crust, prior to subduction. Diamonds hosting the high δ^{18} O garnets have a crustal δ^{13} C 29 signature (-29.6 to -19.4 ‰) paired with δ^{15} N values (-5.6 to +1.3 ‰, with one outlier at +9.9 ‰) 30 implying typically only minor ¹⁵N enrichment. This apparent decoupling of δ^{18} O and δ^{13} C from δ^{15} N 31 indicates diamond formation in an eclogitic substrate where garnet oxygen isotope compositions are 32 inherited from altered low-pressure protoliths, diamond carbon is principally derived from biogenic 33 34 carbonate and organic matter, and diamond nitrogen is variably added by an external, mantle-derived 35 fluid that prompted diamond formation. Of the two remaining garnet inclusions, one has mantle-like $\delta^{18}O$ (+5.4 ‰) and a positive Eu anomaly, suggesting derivation from gabbroic protoliths originally 36 located deep in the oceanic crust where significant alteration by seawater did not occur. Based on 37 mantle-like δ^{13} C (-4.7 ‰) and δ^{15} N (-6.9 ‰), the associated diamond formed from mantle-derived 38 39 fluids/melts. Relative to the other Koidu garnet inclusions, the final garnet has a small majorite component (formation pressure ~ 8 GPa), high Mg# (79.0), elevated Cr# (0.90), and low δ^{18} O (+6.3 ‰), 40 41 suggesting encapsulation during infiltration of slab-derived melts into surrounding asthenospheric 42 peridotite, associated with a high degree of chemical exchange between slab- and mantle-derived43 components.

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45 **Keywords:** Eclogitic diamond; carbon isotope; nitrogen isotope; oxygen isotope; slab-mantle interaction

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47 **1. Introduction**

48 Despite the minor proportion of eclogite within the lithospheric upper mantle (< 1 vol%; Schulze, 1989), eclogite xenoliths provide key information about past geodynamic processes. Subduction and 49 50 associated metamorphism of oceanic crust in the mantle (Helmstaedt and Doig, 1975; MacGregor and 51 Manton, 1986; Jacob, 2004) has been widely accepted as the origin of most, if not all, mantle eclogite 52 xenoliths. Alternate models that invoke direct crystallization from basaltic melts produced from peridotites in the deep mantle (O'Hara and Yoder, 1967; McGetchin and Silver, 1972; Caporuscio and 53 54 Smyth, 1990) are inconsistent with the major and trace element relationships observed in eclogite xenoliths (Aulbach and Arndt, 2019). In addition, Eu anomalies and oxygen isotope signatures in 55 eclogites are unambiguous evidence for a crustal origin. 56

Europium anomalies in the rare earth element (REE) patterns of eclogite whole rocks or their mineral constituents indicate that the formation of their protoliths involved accumulation or fractionation of plagioclase – a low-pressure mineral formed in Earth's crust. Since plagioclase preferentially incorporates Eu from melts relative to other REEs (Weill and Drake, 1973; Bédard, 1994), cumulate gabbros, cumulate-enriched in plagioclase that crystallized in the lower oceanic crust, often have positive Eu anomalies. This results in depletion of Eu in the residual melts, which subsequently may rise into the upper oceanic crust to form basalts with negative Eu anomalies (Philpotts and Schnetzler, 1968). 64 Since some eclogites do not have discernible Eu anomalies (Jacob, 2004; Schmickler et al., 2004), fractionated oxygen isotope compositions are often used to infer a crustal protolith. Peridotitic 65 mantle has a very restricted oxygen isotope composition (δ^{18} O) of +5.5 ± 0.4 ‰ (Mattey et al., 1994; 66 Regier et al., 2018), whereas oceanic crust has a much wider range from 0 to +15 ‰ (Eiler, 2001; 67 Korolev et al., 2018). Oceanic crust with δ^{18} O values above or below the peridotitic mantle range is 68 attributed to low-temperature (< 150 °C) submarine weathering or high-temperature (> 350 °C) 69 70 hydrothermal alteration, respectively. Low-temperature weathering occurs at or near the seawater-basalt interface, while high-temperature alteration occurs in the deeper portion of oceanic crust (Muehlenbachs 71 and Clayton, 1972a,b; Gregory and Taylor, 1981; McCulloch et al., 1981; Alt et al., 1986). Previous 72 studies have shown that the δ^{18} O values inherited from these processes are generally retained by 73 eclogites metamorphosed from subducted oceanic crust (Putlitz et al., 2000; Jacob, 2004; Russell et al., 74 2013). The high δ^{18} O values in eclogites cannot be formed by isotopic fractionation at mantle 75 temperatures from an initial mantle-like value ($+5.5 \pm 0.4$ ‰) (Mattey et al., 1994; Schulze et al., 2013; 76 Regier et al., 2018) nor by mantle metasomatism, as metasomatic fluids with extreme δ^{18} O values are 77 78 not likely to survive unmodified during passage through the dominantly peridotitic mantle (Riches et al., 2016). 79

Garnet is the preferred proxy for the determination of the oxygen isotope composition of eclogites as it is more resistant than clinopyroxene to metasomatic modification and resetting of its isotopic composition (Jacob et al., 1994; Deines and Haggerty, 2000; Schulze et al., 2000; Barth et al., 2001; Korolev et al., 2018). In this study, we analyse garnet inclusions in diamonds from the Koidu kimberlite complex in Sierra Leone for their major element, trace element and stable isotope compositions. Eclogitic diamond formation could be associated with interactions between eclogite and ambient mantle: (1) eclogite-derived carbonated melts may migrate into reducing ambient mantle 87 peridotite and precipitate diamonds (Rohrbach and Schmidt, 2011; Kiseeva et al., 2013; Mikhail et al., 2021); (2) partial melting of eclogite may facilitate *in situ* diamond formation as well as chemical 88 exchange with surrounding peridotite (Smart et al., 2009); (3) mantle-derived fluids/melts may interact 89 90 with eclogitic diamond substrates (Aulbach et al., 2011; Lai et al., 2022). Thus, correlations between δ^{18} O values and major element compositions of garnet inclusions may provide insights into the degree of 91 slab-mantle interaction during diamond formation. Carbon (δ^{13} C) and nitrogen (δ^{15} N) isotope 92 93 compositions of the host diamonds of these garnet inclusions are analysed to identify any mixing of crust- and mantle-derived carbon and nitrogen during this process. Here we present the first coupled 94 δ^{13} C– δ^{15} N– δ^{18} O measurements of diamonds and their inclusions, to gain a better understanding of the 95 96 nature of the interaction between subducted slabs and ambient mantle during diamond formation.

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98 **2. Samples and methods**

99 The Koidu kimberlite complex is situated in the Kono District of eastern Sierra Leone, in the 100 southern part of the West African Craton. Although eclogites are the only type of mantle xenoliths 101 recovered from Koidu (Tompkins and Haggerty, 1984; Hills and Haggerty, 1989; Fung and Haggerty, 102 1995), kimberlite indicator minerals and diamonds from this locality are of both peridotitic and eclogitic 103 paragenesis (Deines and Harris, 1995; Skinner et al., 2004; Harder et al., 2013).

A mineral chemistry study on 105 Koidu diamonds revealed a dominance of eclogitic diamonds (78%), followed by lesser proportions of peridotitic (17%) and mixed paragenesis diamonds (containing co-occurring peridotitic and eclogitic mineral inclusions; 5%) (Lai et al., 2022). Sixteen eclogitic garnet inclusions from 15 different eclogitic diamonds and one mixed paragenesis diamond were selected for the determination of their oxygen isotope composition. The selected garnets cover the entire major element compositional range for all Koidu eclogitic garnet inclusions, and thus are representative of thesample suite.

111 Methods for determining major, minor and trace element compositions of mineral inclusions 112 were previously described in detail (Lai et al., 2022) and are only summarised here. Garnet inclusions were mounted in epoxy and analysed with a CAMECA SX100 electron probe microanalyzer (EPMA) 113 114 for their major and minor element compositions at an accelerating voltage of 20 kV, a beam current of 115 20 nA, and a fully focused beam with a diameter $< 1 \mu m$. Oxide detection limits are typically ≤ 0.02 wt%. Subsequently, trace element compositions of garnet inclusions were determined using a Resonetics 116 117 M-50-LR 193 nm ArF excimer laser ablation system coupled with a Thermo Scientific Element IIXR inductively coupled plasma mass spectrometer (LA-ICP-MS). Samples were ablated with a spot size of 118 23–90 µm at a frequency of 10 Hz and a laser fluence of ~4 J/cm². Calcium contents of garnet inclusions 119 120 determined with EPMA were used as internal standard. Detection limits for trace element concentrations are typically \leq 40 ppb for REE, V, Rb, Sr, Zr, Nb, Ba and Hf, and \leq 1 ppm for Ti and Ni. 121

122 Selected eclogitic garnet inclusions were subsequently co-mounted into a single 25 mm epoxy mount with reference materials S0068 (pyrope garnet from the Gore Mountain area; Bartholomé, 1960) 123 and S0088B (grossular garnet from the Jeffrey mine; Akizuki, 1989). The epoxy mount was polished 124 125 and coated with 25 nm of Au prior to scanning electron microscopy (SEM). Secondary electron (SE) and backscattered electron (BSE) images were obtained using a Zeiss EVO MA15 SEM operating at a 126 voltage of 20 kV and a beam current of 3-4 nA. After imaging, the epoxy mount was coated with 100 127 nm of Au. Oxygen isotope compositions were then determined using a CAMECA IMS-1280 multi-128 collector ion microprobe. Three to six spots were analysed on each garnet inclusion to assess 129 compositional homogeneity. A ¹³³Cs⁺ primary beam with an impact energy of 20 keV and a beam 130 131 current of 2 nA was focused to a beam diameter of approximately 12 µm. Analysis of each spot took

240 s. Oxygen isotope compositions are reported as $\delta^{18}O_{VSMOW}$ [normalized difference of the ${}^{18}O/{}^{16}O$ 132 ratio of the sample relative to that of Vienna Standard Mean Ocean Water (VSMOW), where 133 $^{18}\text{O}/^{16}\text{O}_{\text{VSMOW}} = 2005.20 \ (\pm 0.45) \times 10^{-6} \ (\text{Baertschi}, 1976)].$ Reference materials S0068 (with $\delta^{18}\text{O}_{\text{VSMOW}}$ 134 = +5.72 ‰) and S0088B (with $\delta^{18}O_{VSMOW}$ = +4.13 ‰) were analysed after every four and eight 135 unknowns, respectively, to monitor the instrumental mass fractionation (IMF). The 95% confidence 136 uncertainty estimates for δ^{18} Ovsmow of analysed garnet inclusions average ± 0.26 %, which include 137 138 errors related to within-spot counting statistics, between-spot (geometric) effects, correction for IMF, and matrix effects associated with Ca# (i.e., molar Ca/(Ca+Mg+Fe)). Details of the analytical procedure 139 140 and matrix calibration are outlined in supplementary information and references therein.

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142 **3. Results**

143 **3.1 Major element composition**

The selected garnet inclusions show large variations in CaO (3.6-18.6 wt%), MgO (8.5-20.6 144 wt%), FeO (9.0-18.6 wt%) and Mg# (49.1-79.0), with the garnet (138-7-3) from a mixed paragenesis 145 diamond having the highest CaO and lowest MgO and FeO (Table 1). The garnet (140-2-1) with the 146 147 highest MgO content also has the highest TiO_2 concentration (0.86 vs 0.24–0.48 wt% in other garnets). The high Ti in this garnet does not correspond to its relatively low Na₂O (0.13 wt%). Typically, Na 148 increases with Ti in garnet (Fig. 1), corresponding to the coupled substitution: $M^{2+} + Al^{3+} = Na^+ + Ti^{4+}$ 149 150 (Ringwood and Major, 1971; Bishop et al., 1976, 1978; Grew et al., 2013), however, this positive correlation is not observed for majoritic garnets (Stachel et al., 1998). The concentration of Ti in garnet 151 increases with temperature (Aulbach, 2020) and consequently with pressure, along a conductive 152 geotherm. This agrees with high pressure-high temperature experiments showing that Ti in garnet 153

154 increases with both temperature and pressure (Zhang et al., 2003). This garnet also has relatively low 155 Al_2O_3 (21.1 wt%) compared with the average Al_2O_3 content (22.9 wt%) of other garnets in this study. A negative correlation ($R^2 = 0.81$; n = 16) between Al and Ti cations is observed for garnet inclusions from 156 157 Koidu diamonds (Fig. 2), which is consistent with experimental results documenting decreasing Al and increasing Ti with increasing temperature and pressure (Zhang et al., 2003), following the coupled 158 substitution: $2AI^{3+} = M^{2+} + Ti^{4+}$ (Ackerson et al., 2017). In agreement, garnet 140-2-1 has a minor 159 majorite component of 3.4 mol% (calculated following Locock, 2008). Compared to Si and Al + Cr 160 cation contents in experimental majoritic garnets (Akaogi and Akimoto, 1979; Irifune et al., 1986; 161 Irifune, 1987; Collerson et al., 2010), a pressure of 7.8 GPa (corresponding to a depth of 250 km) (Fig. 3) 162 is estimated for this garnet. Applying different formulations of the majorite-in-garnet barometer 163 (Collerson et al., 2010; Beyer and Frost, 2017) yields equilibration pressures of 7.8-8.2 GPa, 164 corresponding to a depth between 250 and 260 km. This depth range is slightly below the lithosphere-165 asthenosphere boundary of the West African Craton (220 ± 10 km), determined from Koidu 166 clinopyroxene xenocryst geothermobarometry (Smit et al., 2016). Backscattered electron imaging of 167 garnet 140-2-1 shows no exsolution, suggesting that it did not re-equilibrate at shallower depth, and was 168 brought to the surface directly from the uppermost asthenosphere. 169

170 **3.2 Trace element composition**

Koidu eclogitic garnet inclusions are depleted in LREE with subchondritic La, which is a typical feature observed in eclogitic garnets from mantle xenoliths (Jacob, 2004) and included in diamonds (Stachel et al., 2004) worldwide. The garnet (138-7-3) from the mixed paragenesis diamond is distinct in having the highest LREE (La_N ~0.3) and lowest HREE (Lu_N ~6) contents. The remaining garnets have flat MREE_N-HREE_N with Lu_N > 20 (Table 2), and among them, the mildly majoritic garnet 140-2-1 has the highest LREE and lowest HREE. Three garnets (131-5-8, 138-6-2 and 138-12-1) have negative Eu anomalies (Fig. 4a), with [Eu/Eu*]_N = 0.80–0.89, where Eu* = $[Sm_N \times Gd_N]^{0.5}$ (McLennan, 1989). Garnet 138-7-3 has a small positive Eu anomaly ([Eu/Eu*]_N = 1.10), but two additional garnet inclusions from the same diamond have more prominent positive Eu anomalies ([Eu/Eu*]_N = 1.36–1.41). All other garnets in this study lack discernible Eu anomalies (Fig. 4b).

182 **3.3 Oxygen isotope composition**

Backscattered electron images and multiple spot analyses indicate that all garnet inclusions in this study have homogeneous compositions, which excludes the possibility of alteration by fluid/melt infiltrating through cracks in their host diamonds. The internal variability of δ^{18} O for individual grains (average $2\sigma = 0.17$ ‰) is within analytical uncertainty ($2\sigma = ~0.26$ ‰), thus only average δ^{18} O values are reported for each grain (Table 3) and used for discussion in the following section. Oxygen isotope compositions for all analyzed spots are provided in Table S1.

The range of δ^{18} O values of garnet inclusions in this study (+5.4 to +12.1 ‰; Fig. 5) is much 189 greater than that of garnets from the Koidu low-MgO (+4.7 to +6.8 ‰) and high-MgO (+5.1 to +5.7 ‰) 190 eclogite xenoliths (Barth et al., 2001, 2002). Garnet 138-7-3 from the mixed paragenesis diamond has 191 the lowest δ^{18} O value (+5.4 ‰), falling within the mantle range (+5.5 ± 0.4 ‰; Mattey et al., 1994). The 192 mildly majoritic garnet 140-2-1 also has a relatively low δ^{18} O value (+6.3 ‰). The other garnets in this 193 study have δ^{18} O values from +9.9 to +12.1 ‰, with a median value of +11.6 ‰, which is higher than 194 195 that of any other eclogitic garnet inclusion in lithospheric diamonds worldwide, including diamonds from the Damtshaa kimberlites (+4.7 to +8.8 %; Ickert et al., 2013), Argyle lamproite (+6.0 to +8.3 %; 196 Schulze et al., 2013), Finsch kimberlite (+5.7 to +8.0 ‰; Lowry et al., 1999), Siberian placer deposits 197 (+4.8 to +9.6 %; Zedgenizov et al., 2016), Jericho kimberlite (+5.2 to +6.0 %; Smart et al., 2012) and 198 Guaniamo placer deposit (+7.9 to +11.3 %; Schulze et al., 2004) (Fig. 6). 199

A negative correlation between δ^{18} O values and Mg# (molar 100×Mg/(Mg+Fe)) and a positive 200 correlation between δ^{18} O values and Ca# (molar 100×Ca/(Ca+Mg+Fe)) are observed in some garnets 201 (Fig. 7a and b). These correlations are especially obvious for three garnets (138-3-1, 138-5-1 and 138-202 203 11-1) that also show decreasing LREE contents with decreasing Ca# (Fig. 4b). Garnets with the highest δ^{18} O values (~+12 ‰) form a tight cluster in Ca# (0.14–0.18) (Fig. 7b). Except for garnet 138-7-3, all 204 garnets follow the trend ($R^2 = 0.77$; n = 15) of increasing Cr# (molar 100×Cr/(Cr+Al)) with decreasing 205 δ^{18} O values. The mildly majoritic garnet 140-2-1 has the highest Cr# and lowest δ^{18} O along this trend 206 (Fig. 7c). 207

208

209 **4. Discussion**

210 4.1. Origin of Koidu eclogites

Three garnet inclusions (Fig. 4a) with negative Eu anomalies indicate that the protoliths of their host eclogites formed from Eu-depleted residual melts as a result of preceding fractionation of plagioclase. The high δ^{18} O values in these three samples, and almost all other analyzed garnet inclusions (+9.9 to +12.1 ‰; Fig. 5), suggest that their protoliths underwent high degrees of submarine weathering at low temperatures (Muehlenbachs and Clayton, 1972b; Alt et al., 1986; McCulloch et al., 1981). Alteration must have occurred close to the basalt-seawater interface, as δ^{18} O values > +8 ‰ are restricted to altered oceanic crust (AOC) at depths < 300 m (pillow lavas) (Korolev et al., 2018).

The garnet (140-2-1) with a minor majorite component has a δ^{18} O value (+6.3 ‰) that is elevated relative to the mantle range but low compared to other eclogitic garnet inclusions from Koidu. The protolith of this garnet may have been situated slightly deeper in the oceanic crust (e.g., lower portion of the volcanic section), where the water/rock ratio was lower and the alteration temperature was 222 higher, compared to the uppermost portion of the oceanic crust (McCulloch et al., 1981; Alt et al., 1986; Gao et al., 2012). This portion of the oceanic crust then carried the mildly ¹⁸O-enriched signature and 223 subducted through the asthenosphere where pyroxene started dissolving into the garnet structure to form 224 majoritic garnet (Akaogi and Akimoto, 1979; Irifune et al., 1986; Moore and Gurney, 1985). The 225 slightly enriched LREE and depleted HREE contents in this garnet (Fig. 4b) are consistent with an 226 increasing compatibility of LREE and a decreasing compatibility of HREE in garnet as the depth 227 increases (Stachel and Harris, 1997). Alternatively, the relatively low δ^{18} O value and relatively high Cr# 228 in garnet 140-2-1 can be explained by slab-mantle interaction, which will be discussed in the following 229 230 section.

Positive Eu anomalies and low \sum HREE of garnets in diamond 138-7 are consistent with plagioclase accumulation and, consequently, a protolith that originated as cumulate gabbro within the deep oceanic crust (Salisbury and Christensen, 1978; Jacob, 2004; Aulbach and Jacob, 2016). A mantlelike δ^{18} O value (+5.4 ‰) indicates a protolith that was located either too deep in the oceanic crust to be altered by seawater (Alt and Teagle, 2000), or from the intermediate depth zone where the temperature of alteration generates mantle-like δ^{18} O signatures (Schmickler et al., 2004).

4.2. Slab-mantle interaction during diamond formation

Based on their low Cr_2O_3 contents (< 1 wt%; Grütter et al., 2004), all garnet inclusions in this study are classified as eclogitic. It has recently been demonstrated that most majoritic garnets included in diamonds belong to neither a pure eclogitic nor a peridotitic paragenesis, but are intermediate between the two (Kiseeva et al., 2013). An approach based on major element substitutions can better differentiate the parageneses of majoritic garnets, in which peridotitic and pyroxenitic garnets follow the peridotitic majoritic substitution: $2Al^{3+} = Si^{4+} + M^{2+}$, while eclogitic garnets follow the eclogitic majoritic substitution: $M^{2+} + Al^{3+} = Na^+ + Si^{4+}$, with increasing pressure (Kiseeva et al., 2013). Garnet 140-2-1 follows the trend of peridotitic-pyroxenitic majoritic substitution (Fig. S1). Given its high Mg# (79.0)
and elevated Cr# (0.90) relative to the other samples (Table 1) and its low CaO content (4.24 wt%), this
combined evidence suggests derivation from a pyroxenitic lithology.

248 Formation of pyroxenitic inclusions in diamond likely results from the interaction between slabderived melts and ambient mantle peridotite (Aulbach et al., 2002; Kiseeva et al., 2016). This process is 249 250 especially prominent in the sublithospheric mantle, as observed from the high abundance of pyroxenitic 251 majoritic garnets (Kiseeva et al., 2013; Regier et al., 2020). This observation was related to a deep depression in the solidus of carbonated oceanic crust, which intersects most slab geotherms and 252 253 facilitates slab melting at a depth of ~300-700 km (Thomson et al., 2016a). Carbonatite melts derived 254 from subducting slabs can readily migrate into ambient mantle peridotite because of their low dihedral angle with olivine and low viscosity, resulting in efficient elemental and isotopic exchange between melt 255 256 and mantle (Minarik and Watson, 1995; Hammouda and Laporte, 2000; Walter et al., 2008; Thomson et al., 2016b). Thus, the relatively low δ^{18} O value (+6.3 ‰), and high Mg# and Cr# in garnet 140-2-1 (Fig. 257 258 7) likely are a consequence of diamond crystallization during carbonatite-mantle peridotite interaction. 259 Indeed, at the pressure of formation of this majoritic garnet (~8 GPa), ambient mantle is expected to be 260 very reducing, with a calculated oxygen fugacity (fO_2) close to the iron-wüstite (IW) buffer (Frost and 261 McCammon, 2008). Reduction of carbonates from oxidizing slab-derived melt by the reducing metalsaturated ambient mantle is the likely mechanism of diamond formation in this scenario (Rohrbach and 262 Schmidt, 2011). 263

The crude correlations between major element compositions and δ^{18} O for a subset of garnet inclusions in this study (Fig. 7) may be a consequence of seawater alteration in different stratigraphic levels of oceanic crust (Gregory and Taylor, 1981; Aulbach and Jacob, 2016), or kimberlite-like ultramafic metasomatism (Aulbach et al., 2020). Alternatively, these covariations could be associated 268 with elemental and isotopic exchange between eclogitic diamond substrates and surrounding peridotitic lithospheric mantle facilitated by melts (e.g., Smart et al., 2009). In this case, the cluster of garnet 269 inclusions with low Cr# (0.07–0.15) and the highest δ^{18} O values (~+12 ‰) may represent the initial 270 271 compositions of the eclogites, without significant chemical exchange with ambient mantle (Fig. 7). Other garnet inclusions that follow the trends of increasing Mg# and Cr# and decreasing Ca# with 272 decreasing δ^{18} O could then be explained by various degrees of interaction with surrounding peridotites. 273 in which inward transport of peridotitic components into eclogite was facilitated by eclogite-derived 274 melts, and the extent of chemical exchange depends on the distance from the eclogite-peridotite 275 interface (Smart et al., 2009). Trace element partitioning between garnet and pyroxene was likely 276 affected by this interaction, as the three garnets (138-3-1, 138-5-1 and 138-11-1) with an obvious 277 positive correlation between Ca# and δ^{18} O also show decreasing LREE contents with decreasing Ca# 278 (Fig. 4b and 7b). The difference in LREE contents could be attributed to crystal-chemical effects, where 279 LREE preferentially partition into garnets with higher Ca# (Harte and Kirkley, 1997; Aulbach et al., 280 2017). 281

Garnet 138-7-3 has a mantle-like δ^{18} O value (+5.4 ‰) and high Ca# (49.7) (Fig. 7b) typical for a 282 cumulate gabbro protolith (e.g., MacGregor and Manton, 1986; Schulze et al., 2000). This garnet 283 originated from deeper layers within oceanic crust that do not show a signature of slab-mantle 284 interaction, implying that its host diamond likely formed during introduction of external carbon, e.g., 285 through asthenosphere-derived fluids/melts (Lai et al., 2022). Alternatively, this diamond may have 286 287 formed as a result of deserpentinization of altered peridotite in the slab within the diamond stability field, where fluids produced from deserpentinization have mantle-like δ^{18} O. Deserpentinization fluids can be 288 either oxidizing or reducing (Debret and Sverjensky, 2017; Piccoli et al., 2019). Depending on the redox 289 290 states of both the overlying crustal portion and the deserpentinization fluids, diamonds can form from

reduction of CO_2 or CO_3^{2-} , or oxidation of CH₄. Mixing of deserpentinization fluids of different redox states within the slab may also facilitate precipitation of diamonds (Aulbach et al., 2012; Stachel et al., 2022).

4.3. Variation of $\delta^{18}O-\delta^{13}C$: mixing between AOC- and mantle-derived carbon?

Coupled $\delta^{18}O - \delta^{13}C$ measurements of garnet-bearing diamonds worldwide (Argyle, Schulze et al., 295 296 2013; Damtshaa, Ickert et al., 2013; Siberian placers, Zedgenizov et al., 2016; Finsch, Lowry et al., 1999; Jagersfontein, Ickert et al., 2015; Collier and Juina, Burnham et al., 2015) exhibit a negative 297 "correlation", where garnets with δ^{18} O near the mantle value are associated with mantle-like carbon 298 $(\delta^{13}C = -5 \pm 3)$; Cartigny et al., 2014), whilst distinctly ¹⁸O-enriched garnets have ¹³C-depleted host 299 diamonds (Fig. 8). Accordingly, diamonds and their garnet inclusions falling along a trend between 300 these two end-members could be interpreted as mixing products between mantle ($\delta^{13}C \approx -5$ %; $\delta^{18}O \approx$ 301 +5.5 ‰) and AOC (δ^{13} C \approx -30 ‰; δ^{18} O \approx +12 ‰) reservoirs, with variable (O/C)_{Mantle}/(O/C)_{AOC} ratios 302 controlling the curvature of different mixing arrays. 303

However, instead of following a full mixing array, Koidu diamonds and their garnet inclusions 304 lie at opposite ends of the supposed trend, where the majority of samples form a cluster with $\delta^{18}O \ge +9.9$ 305 ‰ and δ^{13} C ≤ -19.4 ‰ (Fig. 8 and Table 3). Since there is no known mantle process that can cause such 306 large ¹⁸O enrichment and ¹³C depletion, this main cluster of Koidu samples represents formation of 307 diamonds and their garnet inclusions in a common surface-derived substrate, with diamond carbon and 308 garnet oxygen both being sourced from uppermost AOC (Ickert et al., 2013) containing biogenic 309 carbonate ± organic matter (Milledge et al., 1983; Kirkley et al., 1991; Li et al., 2019). This tight 310 clustering suggests that with respect to oxygen and carbon, the contribution of a mantle-derived 311 component, if present at all, was minor. In subtle contrast, the observed negative correlation between 312 Cr# and δ^{18} O of garnets spanning a δ^{18} O range of +9.9 to +11.3 ‰ may be attributed to eclogite-313

314 peridotite interaction during diamond formation, where the eclogite-derived component dominates the oxygen isotope signature but the small mantle contribution levers the δ^{18} O value downwards (Fig. 7c). 315 In this scenario, the associated dilution of AOC-derived carbon signatures by mantle carbon however, 316 was negligible, due to elevated carbon concentrations in AOC close to the seawater interface relative to 317 carbon-poor ambient mantle. The very low and varied δ^{13} C values of the host diamonds from this cluster 318 (-29.6 to -19.4 ‰) thus directly reflect the variable carbon isotope composition of subducted AOC, with 319 e.g., biogenic carbonate having δ^{13} C as low as -24 ‰ (Li et al., 2019) and organic matter having δ^{13} C of 320 -26 ± 7 ‰ (Schidlowski, 2001). Possible devolatilization of carbonate from AOC during subduction 321 may decrease δ^{13} C values further (Shieh and Taylor, 1969; Li et al., 2019). Although four of the Koidu 322 diamonds containing high δ^{18} O garnets (138-5, 138-9, 143-2 and 146-2) have rim zones with mantle-like 323 δ^{13} C values (Fig. 8), the garnet inclusions were all located in the centres of the diamonds and thus are 324 not related to the rim zones. The abrupt changes in diamond carbon isotope composition from cores (-325 26.5 to -24.0 ‰) to rims (-7.0 to -6.0 ‰) document distinct stages of diamond formation involving 326 carbon from difference sources, with the rim zones being precipitated from mantle-derived fluids/melts 327 328 (Lai et al., 2022).

In contrast to the input of AOC-derived carbon and oxygen required for some Koidu garnet-329 diamond associations, the mantle-like δ^{13} C value (-4.7 ‰) of the diamond hosting garnet 138-7-3 can be 330 explained by a low carbon content in the gabbroic layers of subducting slabs (Li et al., 2019), which 331 requires externally sourced, mantle-derived carbon for the formation of diamonds. However, the scarcity 332 of diamond in eclogites derived from gabbroic layers implies that the available volume of the required 333 mantle-derived fluids/melts was either small or that the redox conditions in the subducted gabbroic 334 substrate were not suitable for diamond formation from the carbon species contained in the mantle-335 336 derived fluids/melts.

The host diamond of majoritic garnet 140-2-1 has elevated δ^{13} C values ranging from -4.2 to 0 % 337 (inner zone: $\delta^{13}C = -4.2$ %); outer zone: $\delta^{13}C = -1.8$ to 0 %); average $\delta^{13}C = -2.6$ %); Fig. 9). Similar 338 mildly ¹³C-enriched carbon isotope compositions were observed on the West African Craton for 339 majoritic garnet-included diamonds from Kankan ($\delta^{13}C = -3.1$ to +0.9 %) and were attributed to 340 diamond formation from subducted normal marine carbonates (Stachel et al., 2002). As discussed in 341 section 4.2, the majoritic garnet with a δ^{18} O value of +6.3 ‰ and its ¹³C-enriched host diamond could 342 have crystallized during the interaction of a slab-derived carbonatitic melt and ambient mantle. The 343 carbon isotope signature of the outer zone of this diamond then may be attributed to predominantly slab-344 derived carbon (normal marine carbonates; $\delta^{13}C \approx 0$ %). The mantle-like $\delta^{13}C$ value in the inner zone 345 could be explained by formation of diamond in the peridotitic substrate from mantle-derived carbon 346 prior to or at the initial stages of infiltration of slab-derived melt. This was followed by crystallization of 347 the overgrowth layer (outer zone) along with the majoritic garnet inclusion during melt-peridotite 348 interaction. The drop in nitrogen concentrations and sharp decrease in cathodoluminescence response 349 from the inner ([N] = 51-84 at.ppm) to the outer zone ([N] = 0.6-0.7 at.ppm) (Fig. 9) of this diamond 350 351 clearly documents distinct growth episodes.

4.4. Variation of $\delta^{18}O-\delta^{15}N$: mixing between AOC- and mantle-derived nitrogen?

Through its mantle-like δ^{13} C (-4.7 ‰) and δ^{15} N (-6.9 ‰) values, diamond 138-7, containing a gabbroic garnet inclusion, demonstrates a coupled origin of carbon and nitrogen, likely from asthenosphere-derived fluids/melts (Lai et al., 2022) or from fluids derived from deserpentinization processes (Aulbach et al., 2012). The association of mantle-like δ^{15} N and δ^{18} O (+5.4 ‰) values implies that the gabbroic protolith did not significantly interact with hydrothermal fluids during the seafloor alteration stage (Alt and Teagle, 2000), and thus, in the absence of low-temperature clay formation – as invoked, for example by Li et al. (2019) to explain highly positive δ^{15} N in some diamonds, did not incorporate significant nitrogen with a characteristic ¹⁵N-enriched signature at the site of future diamond
 formation.

362 Other host diamonds in this study with nitrogen contents sufficiently high for isotopic analysis have $\delta^{15}N$ values between -5.6 and +1.3 ‰, except for one outlier with $\delta^{15}N = +9.9$ ‰ (Table 3). The 363 small portion of positive δ^{15} N values in Koidu diamonds is attributed to a contribution of nitrogen from 364 365 clay minerals formed by low-temperature (< 100 °C) alteration of oceanic crust (Busigny et al., 2005; Bebout et al., 2018; Li et al., 2019; Li and Li, 2022). During subduction, devolatilization of nitrogen 366 from AOC further increases δ^{15} N values in subducted materials (Bebout and Fogel, 1992; Cartigny et al., 367 2014), resulting in diverse δ^{15} N values in AOC. Some Koidu diamonds are associated with convecting 368 mantle-like δ^{15} N values (-5.6 to -2.0 ‰), but with clearly subducted oxygen (δ^{18} O = +10.3 to +11.8 ‰) 369 and carbon ($\delta^{13}C = -26.5$ to -22.0 %) isotope signatures. This suggests the possibility of independent 370 371 sources for (1) carbon and oxygen and (2) nitrogen during the diamond-forming process, with carbon and oxygen originating locally from subducted AOC while nitrogen was largely derived from the mantle. 372 This decoupling of carbon and nitrogen isotope signatures is indeed relatively common in diamonds 373 374 worldwide (Cartigny et al., 2014; Mikhail et al, 2014).

To investigate the extent of mixing of nitrogen derived from AOC and the mantle during 375 376 diamond precipitation, we model mixing arrays using different oxygen/nitrogen (O/N) ratios of AOC and the mantle (see supplementary information for details of modeling methods). Mixing models 377 involve two major reservoirs as the endmembers: the mantle ($\delta^{15}N = -5.0$ %; $\delta^{18}O = +5.5$ %) and AOC 378 (represented by diamond 131-2 with $\delta^{15}N = +9.9$ % and $\delta^{18}O = +12.1$ %). Curvature of the mixing 379 arrays is controlled by $n = (O/N)_{AOC}/(O/N)_{mantle}$ ratios. Note that the oxygen contents of eclogite and 380 peridotite are similar (Cartigny et al., 2014), thus the effect of changing oxygen content on the O/N ratio 381 382 is negligible and the major parameter that controls the curvature of the mixing model is the nitrogen

content of the two interacting reservoirs. Koidu diamonds are best fitted by the mixing curves with n > 1383 (e.g., n = 10, 25 and 70 shown in Fig. 10). These mixing curves require a mantle reservoir with higher 384 nitrogen content compared to AOC, indicating that the contribution of nitrogen from the mantle during 385 386 diamond formation is high (mantle-derived nitrogen > 60%; Fig. 10). This could be explained by diamond formation in eclogitic substrates facilitated by a mantle-derived fluid/melt pulse, in which the 387 mantle-derived fluid/melt carried a comparatively high concentration of nitrogen but negligible carbon, 388 thus adding a mantle-like δ^{15} N signature to the diamonds without affecting their crustal δ^{13} C signature. 389 Crustal nitrogen remains stable in slabs as NH⁴⁺ in potassium-bearing minerals and transfers to the deep 390 mantle only in cold subduction zones; in warm subduction zones, however, there is substantial nitrogen 391 loss during slab devolatilization and, thus, the amount of crustal nitrogen left in the downgoing slab for 392 subsequent diamond formation becomes negligible (Labidi, 2022). 393

394 An alternative explanation is provided by observations of oxygen and nitrogen isotope compositions of the altered basaltic sections recovered at Ocean Drilling Program Site 801, which are 395 characterized by high δ^{18} O (+8.7 to +25.7 %; Alt, 2003) and a wide range of δ^{15} N (-11.6 to +1.2 %; Li 396 et al., 2007) values. The isotope compositions of the super-composite (constructed to represent the bulk 397 composition of the upper oceanic crust; Alt, 2003) at Site 801 show mantle-like δ^{15} N values (-5.4 %; Li 398 et al., 2007) with very high δ^{18} O values (+12.0 %; Alt, 2003). This suggests the possibility that mantle-399 400 like δ^{15} N values of some Koidu diamonds may still derive from an AOC source and as such, their nitrogen isotope compositions are not "decoupled" from the carbon and oxygen isotope signatures 401 402 recorded by the diamond-garnet association at Koidu.

403

404 **5.** Conclusions

405 The coupled $\delta^{13}C - \delta^{15}N - \delta^{18}O$ signatures and elemental compositions of diamonds and their 406 eclogitic inclusions from the Koidu mine suggests three possible modes of diamond formation during 407 slab-mantle interaction:

408 (1) In the gabbroic layer of oceanic crust where the concentration of crustal carbon is low, diamond 409 formation relies on carbon sourced from mantle-derived fluids/melts, thus the single Koidu 410 diamond formed in this layer has mantle-like carbon and nitrogen isotope compositions ($\delta^{13}C = -$ 411 4.7 ‰; $\delta^{15}N = -6.9$ ‰). The protolith of this diamond substrate was not significantly altered by 412 seawater and consequently, the gabbroic garnet inclusion has an unperturbed oxygen isotope 413 composition ($\delta^{18}O = +5.4$ ‰).

(2) In the uppermost basaltic layer of altered oceanic crust, the concentration of carbon is high and 414 dominates the carbon budget for diamond formation. Diamonds with strongly ¹³C-depleted 415 carbon isotope compositions (δ^{13} C = -29.6 to -19.4 ‰) document carbon derived from biogenic 416 carbonate and/or organic matter, with a negligible contribution of normal marine carbonates. 417 Nitrogen isotope compositions of this diamond population fall in an only mildly ¹⁵N enriched 418 range between -5.6 and +1.3 % (one outlier at +9.9 %). This suggests that either the nitrogen 419 budget during diamond formation was dominated by mantle-derived nitrogen, arriving from an 420 external, mantle-hosted fluid source, or the subducted AOC substrate for Koidu diamonds had an 421 unusual nitrogen isotope composition (without ¹⁵N enrichment), similar to that observed in 422 oceanic crust at ODP site 801. The basaltic protoliths that formed the host eclogites to this suite 423 424 of diamonds were intensely altered by seawater at low temperatures prior to subduction, imparting strongly ¹⁸O-enriched isotopic compositions ($\delta^{18}O = +9.9$ to +12.1 ‰) on garnet 425 inclusions derived from this layer. If not related to stratigraphic position of the protoliths within 426 the upper oceanic crust or kimberlite-like ultramafic metasomatism, covariations between major 427

element and oxygen isotope compositions of garnet inclusions suggest chemical exchange
between eclogite and ambient mantle peridotite coeval with diamond formation. This exchange
may have been facilitated by melt generated during partial melting of eclogite, in which the
degree of chemical exchange depends on the distance of garnets from the eclogite-peridotite
interface, i.e., the distance of melt-assisted inward diffusion of elements from peridotite.

(3) One Koidu diamond containing a mildly majoritic garnet inclusion indicates that slab-derived 433 fluids/melts infiltrated ambient mantle peridotite and precipitated pyroxenitic diamonds. The 434 core and rim structure of the precipitated diamond suggests that the core formed from 435 predominantly mantle-derived carbon (core zone $\delta^{13}C = -4.2$ %) and the rim from slab-derived 436 crustal carbon (rim zone δ^{13} C = -1.8 to 0 ‰). Peridotite can efficiently buffer the oxygen isotope 437 composition of infiltrating slab-derived fluids/melts, driving the δ^{18} O value of the garnet 438 inclusion (+6.3) towards the mantle range compared with the much higher values observed in 439 other garnets within Koidu diamonds. The high Mg#, elevated Cr# and low Ca# of the garnet 440 inclusion are all consistent with a high degree of buffering of slab-derived components during 441 442 interaction with the mantle.

443

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451	Appendix A. Supplementary Material
452	Details of SIMS analysis and mixing model are provided in the methodology section in supplementary
453	material. Oxygen isotope compositions for all analyzed spots are shown in Table S1. Trends of
454	peridotitic-pyroxenitic and eclogitic majoritic substitution are shown in Fig. S1.
455	
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Sample	131-2-2	131-5-8	137-1-6	138-2-1	138-3-1	138-4-1	138-5-1	138-6-2
SiO_2	40.53	40.29	40.22	41.02	40.91	40.38	41.26	40.06
TiO ₂	0.36	0.36	0.37	0.47	0.36	0.39	0.39	0.24
Al_2O_3	22.68	22.67	21.99	23.08	23.48	22.94	23.27	22.72
Cr_2O_3	0.02	0.05	0.04	0.15	0.07	0.04	0.08	0.11
FeO ^{total}	17.98	16.58	18.58	13.52	17.04	18.32	15.27	15.94
NiO	n.d.	n.d.	n.d.	n.d.	n.d.	n.d.	n.d.	0.01
MnO	0.24	0.38	0.26	0.26	0.25	0.24	0.25	0.29
MgO	12.11	10.13	11.27	15.01	13.04	11.60	15.66	12.19
CaO	6.03	9.59	6.22	6.03	5.91	6.23	4.32	7.24
Na ₂ O	0.17	0.16	0.16	0.17	0.16	0.18	0.13	0.12
Total	Total 100.11 100.23		99.09	99.71	101.22	100.31	100.63	98.92
Sample	138-7-3	138-9-6	138-11-1	138-12-1	140-2-1	143-2-3	144-1-3	146-2-1
SiO ₂	41.00	40.70	41.38	40.48	42.56	40.59	40.38	40.56
TiO ₂	0.47	0.38	0.48	0.25	0.86	0.29	0.32	0.33
Al_2O_3	22.98	22.38	23.13	22.88	21.09	22.85	22.94	23.38
Cr_2O_3	0.02	0.04	0.17	0.04	0.28	0.05	0.05	0.04
FeO ^{total}	8.99	17.57	14.56	16.64	9.79	17.45	18.00	17.41
NiO	n.d.	n.d.	n.d.	n.d.	0.01	n.d.	n.d.	n.d.
MnO	0.16	0.22	0.26	0.41	0.27	0.24	0.23	0.35
MgO	8.47	12.26	16.83	9.00	20.61	12.33	11.67	13.06
CaO	18.55	6.21	3.57	11.08	4.24	6.29	6.58	5.39
Na ₂ O	0.18	0.16	0.14	0.13	0.13	0.14	0.14	0.14
Total	100.82	99.93	100.52	100.91	99.82	100.23	100.32	100.67

Table 1. Major element compositions (wt%) of garnet inclusions in 16 Koidu diamonds.

n.d. = not determined.

<u>±</u>

Sample	131-2-2	131-5-8	137-1-6	138-2-1	138-3-1	138-4-1	138-5-1	138-6-2
La	0.005	0.010	0.010	0.021	0.023	0.010	0.009	0.009
Ce	0.149	0.339	0.198	0.428	0.392	0.295	0.148	0.091
Pr	0.104	0.227	0.134	0.159	0.123	0.129	0.084	0.070
Nd	1.347	3.115	1.497	1.483	1.392	1.470	1.067	1.316
Sm	3.100	2.914	3.017	2.672	1.976	3.260	1.362	2.311
Eu	2.084	1.118	1.593	1.322	1.120	1.978	0.724	0.977
Gd	8.880	5.080	9.180	6.060	5.840	9.100	4.360	6.000
Tb	1.688	0.967	1.373	1.087	1.164	1.887	1.061	1.002
Dy	11.420	6.810	9.460	7.270	8.030	13.090	8.560	7.360
Но	2.404	1.504	2.473	1.529	1.635	2.896	1.976	1.875
Er	7.170	4.651	8.300	4.584	4.893	8.860	5.990	6.250
Tm	1.038	0.661	1.015	0.655	0.697	1.255	0.855	0.807
Yb	7.320	4.736	7.050	4.659	4.882	8.640	5.805	5.510
Lu	1.129	0.726	1.093	0.707	0.737	1.337	0.835	0.836
Sample	138-7-3	138-9-6	138-11-1	138-12-1	140-2-1	143-2-3	144-1-3	146-2-1
La	0.062	0.004	0.002	0.022	0.050	0.037	0.007	0.009
Ce	0.559	0.105	0.000	0.812	0.374	0.418	0.291	0.249
Pr	0.213	0.077	0.061	0.623	0 100	0.100	0.220	0 172
Nd				0.025	0.108	0.192	0.239	0.175
110	1.910	1.171	0.784	8.000	0.108 0.934	0.192 2.040	0.239 2.378	0.173 1.414
Sm	1.910 1.314	1.171 3.324	0.784 1.218	8.000 5.620	0.108 0.934 0.911	0.192 2.040 2.700	0.239 2.378 3.254	0.173 1.414 2.044
Sm Eu	1.910 1.314 0.617	1.171 3.324 1.563	0.784 1.218 0.656	8.000 5.620 1.931	0.108 0.934 0.911 0.475	0.192 2.040 2.700 1.660	0.239 2.378 3.254 1.579	0.173 1.414 2.044 1.227
Sm Eu Gd	1.910 1.314 0.617 2.240	1.171 3.324 1.563 8.440	0.784 1.218 0.656 3.603	8.000 5.620 1.931 8.570	0.108 0.934 0.911 0.475 2.441	0.192 2.040 2.700 1.660 7.520	0.239 2.378 3.254 1.579 8.280	0.173 1.414 2.044 1.227 6.760
Sm Eu Gd Tb	1.910 1.314 0.617 2.240 0.321	1.171 3.324 1.563 8.440 1.144	0.784 1.218 0.656 3.603 0.810	8.000 5.620 1.931 8.570 1.571	0.108 0.934 0.911 0.475 2.441 0.576	0.192 2.040 2.700 1.660 7.520 1.622	0.239 2.378 3.254 1.579 8.280 1.176	0.173 1.414 2.044 1.227 6.760 1.097
Sm Eu Gd Tb Dy	1.910 1.314 0.617 2.240 0.321 2.138	1.171 3.324 1.563 8.440 1.144 7.630	0.784 1.218 0.656 3.603 0.810 6.600	8.000 5.620 1.931 8.570 1.571 11.250	0.108 0.934 0.911 0.475 2.441 0.576 4.732	0.192 2.040 2.700 1.660 7.520 1.622 11.680	0.239 2.378 3.254 1.579 8.280 1.176 7.850	0.173 1.414 2.044 1.227 6.760 1.097 8.460
Sm Eu Gd Tb Dy Ho	1.910 1.314 0.617 2.240 0.321 2.138 0.540	1.171 3.324 1.563 8.440 1.144 7.630 1.898	0.784 1.218 0.656 3.603 0.810 6.600 1.578	0.023 8.000 5.620 1.931 8.570 1.571 11.250 2.477	0.108 0.934 0.911 0.475 2.441 0.576 4.732 1.128	0.192 2.040 2.700 1.660 7.520 1.622 11.680 2.610	0.239 2.378 3.254 1.579 8.280 1.176 7.850 2.048	0.173 1.414 2.044 1.227 6.760 1.097 8.460 2.246
Sm Eu Gd Tb Dy Ho Er	1.910 1.314 0.617 2.240 0.321 2.138 0.540 1.646	1.171 3.324 1.563 8.440 1.144 7.630 1.898 6.240	0.784 1.218 0.656 3.603 0.810 6.600 1.578 5.000	0.023 8.000 5.620 1.931 8.570 1.571 11.250 2.477 7.370	0.108 0.934 0.911 0.475 2.441 0.576 4.732 1.128 3.634	0.192 2.040 2.700 1.660 7.520 1.622 11.680 2.610 8.170	0.239 2.378 3.254 1.579 8.280 1.176 7.850 2.048 6.750	0.173 1.414 2.044 1.227 6.760 1.097 8.460 2.246 7.850
Sm Eu Gd Tb Dy Ho Er Tm	1.910 1.314 0.617 2.240 0.321 2.138 0.540 1.646 0.182	1.171 3.324 1.563 8.440 1.144 7.630 1.898 6.240 0.775	0.784 1.218 0.656 3.603 0.810 6.600 1.578 5.000 0.756	0.023 8.000 5.620 1.931 8.570 1.571 11.250 2.477 7.370 1.054	0.108 0.934 0.911 0.475 2.441 0.576 4.732 1.128 3.634 0.535	0.192 2.040 2.700 1.660 7.520 1.622 11.680 2.610 8.170 1.217	0.239 2.378 3.254 1.579 8.280 1.176 7.850 2.048 6.750 0.836	0.173 1.414 2.044 1.227 6.760 1.097 8.460 2.246 7.850 0.988
Sm Eu Gd Tb Dy Ho Er Tm Yb	$ 1.910 \\ 1.314 \\ 0.617 \\ 2.240 \\ 0.321 \\ 2.138 \\ 0.540 \\ 1.646 \\ 0.182 \\ 1.116 $	1.171 3.324 1.563 8.440 1.144 7.630 1.898 6.240 0.775 5.420	0.784 1.218 0.656 3.603 0.810 6.600 1.578 5.000 0.756 5.470	0.023 8.000 5.620 1.931 8.570 1.571 11.250 2.477 7.370 1.054 7.640	0.108 0.934 0.911 0.475 2.441 0.576 4.732 1.128 3.634 0.535 3.835	0.192 2.040 2.700 1.660 7.520 1.622 11.680 2.610 8.170 1.217 8.790	0.239 2.378 3.254 1.579 8.280 1.176 7.850 2.048 6.750 0.836 6.070	0.173 1.414 2.044 1.227 6.760 1.097 8.460 2.246 7.850 0.988 6.980

Table 2. Trace element concentrations (ppm) of garnet inclusions in 16 Koidu diamonds.

Sample	δ^{18} Ovsmow (‰)	2σ (‰)	$\delta^{13}C_{VPDB}$ (‰)	2σ (‰)	$\delta^{15}N_{Air}$ (‰)	2σ (‰)	[N] (at.ppm)	2σ (at.ppm)
131-2-2	12.1	0.3	-19.4	0.1	9.9	2.6	57	2
131-5-8	10.8	0.3	-25.4	0.2	1.1	3.2	41	2
137-1-6	11.8	0.3	-27.9	0.1	0.6	3.5	33	2
138-2-1	11.8	0.3	-28.3	0.1			21	1
138-3-1	10.8	0.3	-26.8	0.1	-0.2	0.9	552	17
138-4-1	11.8	0.3	-26.3	0.1	-2.6	1.3	267	9
138-5-1 (rim)	10.3	0.3	-7.0	0.1	-5.8	0.8	968	30
138-5-1 (core)	10.5	0.5	-26.5	0.1	-5.6	1.2	290	10
138-6-2	11.1	0.3	-22.0	0.1	-2.0	2.5	55	2
138-7-3	5.4	0.3	-4.7	0.1	-6.9	1.2	515	16
138-9-6 (rim)	12.0	0.3	-6.9	0.1	-5.5	0.8	1027	31
138-9-6 (core)	12.0	0.5	-24.0	0.1	0.2	1.8	71	2
138-11-1	9.9	0.3	-29.6	0.1	1.3	1.2	286	10
138-12-1	11.3	0.3	-28.6	0.1	-1.3	0.9	252	8
140-2-1	6.3	0.3	-2.6	0.1			34	2
143-2-3 (rim)	11.2	0.3	-6.0	0.2	-4.5	0.8	1070	33
143-2-3 (core)	11.5	.5 0.5	-24.9	0.1	-2.9	1.0	133	4
144-1-3	11.9	0.3	-25.7	0.1			21	1
146-2-1 (rim)	11.0	0.2	-7.0	0.1	-4.8	0.9	680	21
146-2-1 (core)	11.9	0.5	-26.2	0.1		—	16	1

Table 3. Average δ^{18} O values and nitrogen concentrations, δ^{13} C and δ^{15} N values of garnet inclusions and their host diamonds, respectively.

















(a)



(b)



(c)









Fig. 1. Ti versus Na (apfu) for garnet inclusions in Koidu diamonds. Except for one garnet with a minor majorite component, garnets show a crude positive correlation ($R^2 = 0.33$) between Ti and Na.

Fig. 2. Al versus Ti (apfu) in garnet inclusions in Koidu diamonds, based on a formula unit containing twelve oxygen atoms. Al and Ti are strongly negatively correlated ($R^2 = 0.81$). The outlier with high Ti content is the mildly majoritic garnet.

Fig. 3. Covariation between Al+Cr and Si in garnets, based on a formula unit containing twelve oxygen atoms. Al+Cr decrease with increasing Si, indicating increasing pressure. Symbol colour represents the depths (km) corresponding to equilibration pressures reported in literature: triangles are data from high pressure experiments conducted by Akaogi and Akimoto (1979); squares are data from high pressure experiments conducted by Irifune et al. (1986) and Irifune (1987); circles are data compiled by Collerson et al. (2010) for garnet inclusions in diamonds worldwide. Black diamond symbols represent the chemical compositions of garnet inclusions in this study. The position of the dashed line, indicating a depth of 220 km (lithosphere-asthenosphere boundary of the West African Craton; Smit et al., 2016), is estimated from the garnet inclusions compiled by Collerson et al. (2010). Estimated pressures for garnet 140-2-1 is 7.8–8.2 GPa (Collerson et al., 2010; Beyer and Frost, 2017), corresponding to a depth between 250 and 260 km.

Fig. 4. REE concentrations in eclogitic garnets from Koidu diamonds normalized to CI-chondrite (McDonough and Sun, 1995). (a) Three garnet inclusions with discernible negative Eu anomalies. (b) Other garnet inclusions lacking obvious Eu anomalies. Garnets from diamond 138-7 (black) are different from other garnets in this study through their negative MREE_N-HREE_N slopes with Lu_N < 10. Two garnets (black dashed lines) from diamond 138-7 have discernible positive Eu anomalies. Inset shows three garnets (138-3-1, 138-5-1 and 138-11-1) that follow a clear trend of decreasing Ca# with decreasing δ^{18} O values (see Fig. 7b).

Fig. 5. Distribution of δ^{18} O values in garnet inclusions from Koidu diamonds. The bin size for the histogram is 0.25 ‰. The region shaded in grey indicates the canonical δ^{18} O range of the mantle (+5.5 ± 0.4 ‰; Mattey et al., 1994). The δ^{18} O values of garnets in this study range from +5.4 to +12.1 ‰, with a mode of +11.8 ‰. Gabbroic garnet 138-7-3 has the lowest δ^{18} O value (+5.4 ‰), which falls within the mantle range.

Fig. 6. Oxygen isotope compositions (δ^{18} O) of eclogitic garnet inclusions in diamonds from Koidu (this study), Damtshaa (Ickert et al., 2013), Argyle (Schulze et al., 2013), Finsch (Lowry et al., 1999), Siberian placer (Zedgenizov et al., 2016), Jericho (Smart et al., 2012) and the Guaniamo placer deposit (δ^{18} O values of multiple garnet inclusions in one single diamond; Schulze et al., 2004). The grey shaded region indicates the canonical δ^{18} O range of the mantle (+5.5 ± 0.4 ‰; Mattey et al., 1994).

Fig. 7. Covariations of oxygen isotope composition (δ^{18} O) and molar Mg# (a), Ca# (b) and Cr# (c) in Koidu garnet inclusions (the gabbroic garnet is shown in blue, other garnets are shown in orange). Three garnets (138-3-1, 138-5-1 and 138-11-1; grey background) with clear correlations between δ^{18} O and major element compositions are shown with symbols following Fig. 4b. The region shaded in grey indicates the canonical δ^{18} O range of the mantle (+5.5 ± 0.4 ‰; Mattey et al., 1994). Garnets show a well-defined negative correlation (R² = 0.77) between Cr# and δ^{18} O values.

Fig. 8. Oxygen isotope composition (δ^{18} O) of garnet inclusions and carbon isotope composition (δ^{13} C) of their host diamonds in this study. Average δ^{18} O and δ^{13} C values of multiple analyzed spots of individual garnet grains and diamond fragments are shown. Four diamonds (138-5, 138-9, 143-2 and 146-2) have a core-rim structure and are linked by dashed tie-lines, with δ^{13} C values of the rims being indicated by white squares and core-zone values by yellow squares. Error bars reflect total uncertainties (95% confidence level) for the δ^{18} O and δ^{13} C values of each sample and for δ^{13} C are smaller than the symbol size. Also shown are δ^{13} C and δ^{18} O values of lithospheric (black)/sublithospheric (red) diamonds and their garnet inclusions from Argyle (Schulze et al., 2013), Damtshaa (Ickert et al., 2013), Siberian placers (Zedgenizov et al., 2016), Finsch (Lowry et al., 1999), Jagersfontein (Ickert et al., 2015) and Collier/Juina (Burnham et al., 2015).

Fig. 9. Cathodoluminescence image of diamond 140-2-1 showing inner and outer growth zones with distinct δ^{13} C values (orange; ‰) and nitrogen concentrations (green; at.ppm).

Fig. 10. Oxygen isotope composition (δ^{18} O) of eclogitic garnet inclusions and nitrogen isotope composition (δ^{15} N) of their host diamonds in this study. Average δ^{18} O and δ^{15} N values of multiple analyzed spots of individual garnet grains and diamond fragments are shown for each sample. Four diamonds (138-5, 138-9, 143-2 and 146-2) have a core-rim structure and are linked by dotted tie-lines, with the compositions of the rims indicated by white squares and core analyses by yellow squares. Error bars reflect total uncertainties (95% confidence level) for the δ^{18} O and δ^{15} N values of each sample. Solid lines indicate mixing arrays calculated using different (O/N)_{AOC}/(O/N)_{mantle} ratios denoted by n: (light grey) n = 10; (dark grey) n = 25; and (black) n = 70. Red dashed lines indicate the percentages of nitrogen contributed by the mantle.